



Systematic Uncertainties in Long Baseline Neutrino Oscillation Measurements



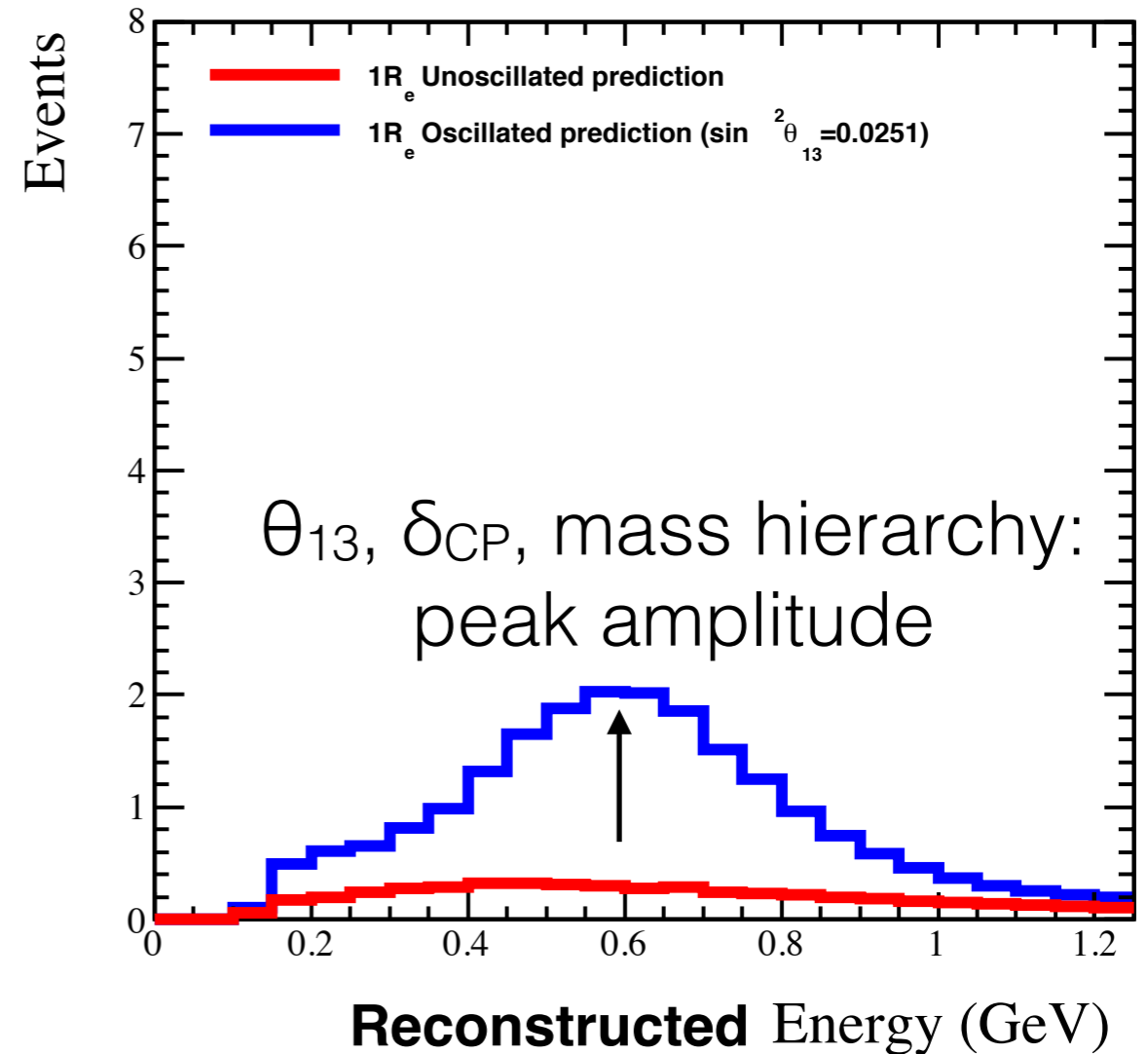
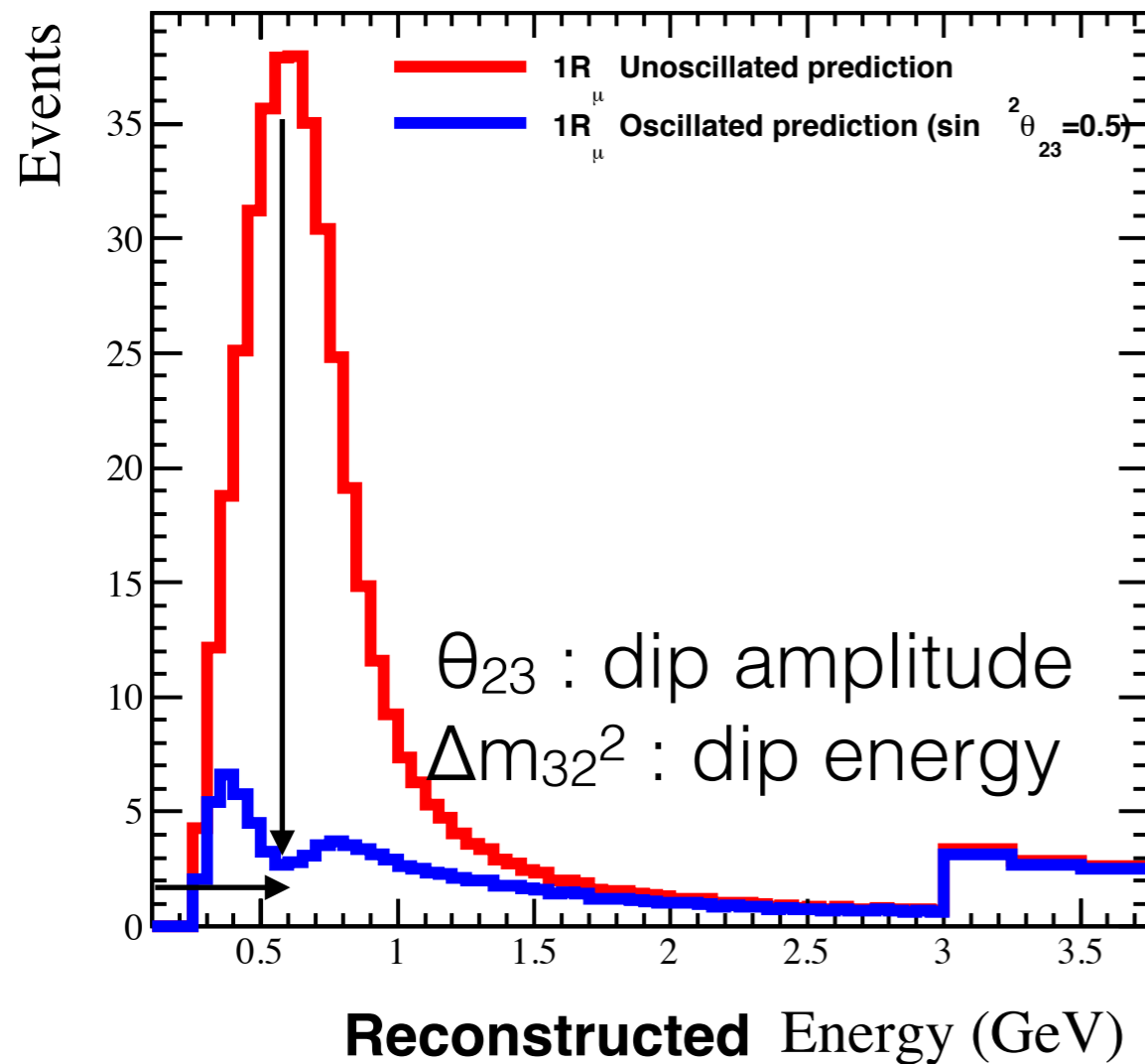
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David Hadley
21st December 2017
Prospects in Neutrino Physics, NuPhys2017

What is measured?

ν_μ disappearance

ν_e appearance



Measurement precision limited by:

- Statistics
- Neutrino energy reconstruction
- Knowledge of unoscillated spectrum and background contamination

Where do systematic uncertainties enter?

$$R(\vec{X}_{\text{reco}}) = \int \Phi(E_\nu) \times \sigma(E_\nu, \vec{X}_{\text{true}}) \times \varepsilon(\vec{X}_{\text{true}}, \vec{X}_{\text{reco}}) \times P(E_\nu, \vec{\theta}) dE_\nu d\vec{X}_{\text{true}}$$

Neutrino
Flux

Neutrino nucleus
interaction model

Detector efficiency
and resolution

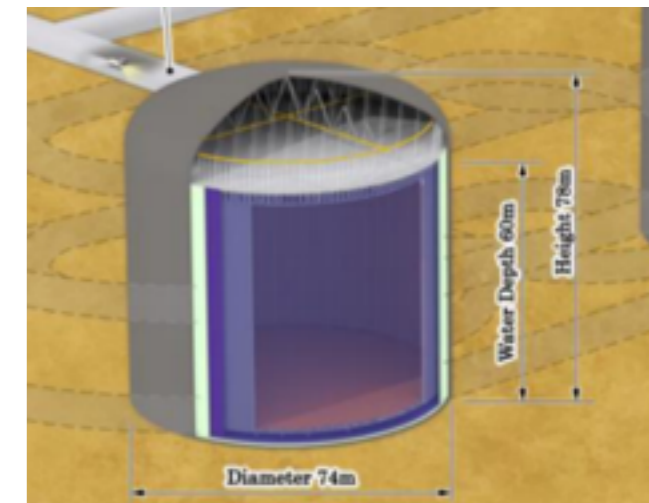
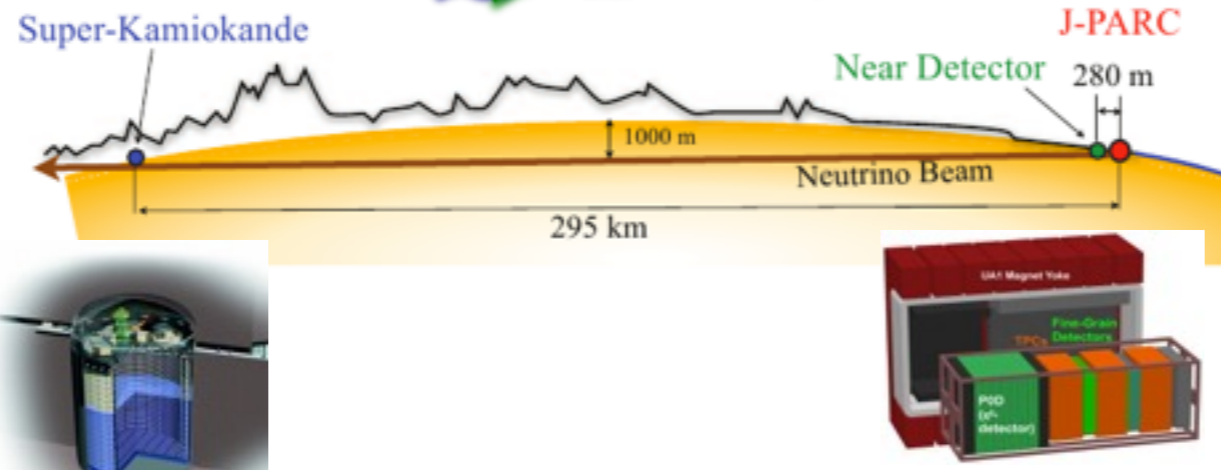
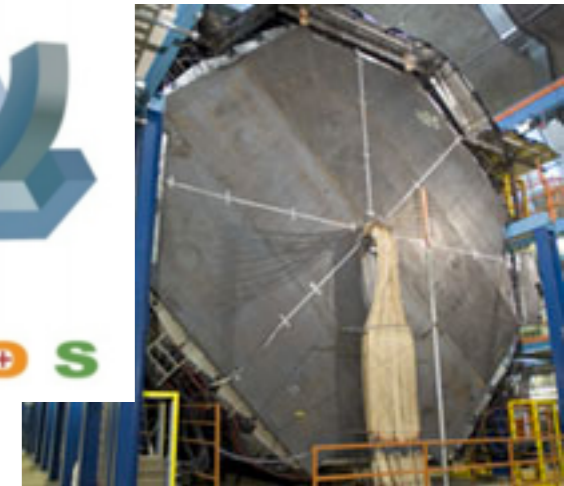
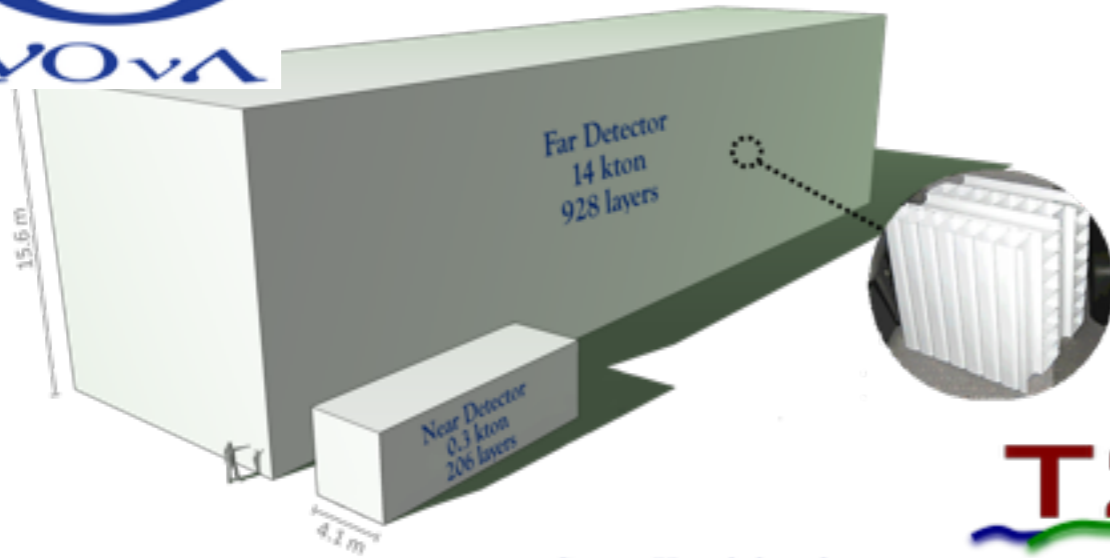
Oscillation
Probability

Use measurements at near detector to constrain $(\Phi \times \sigma \times \varepsilon)$

Cancellation of uncertainties is not perfect as no oscillation at the near detector and Φ and ε may differ

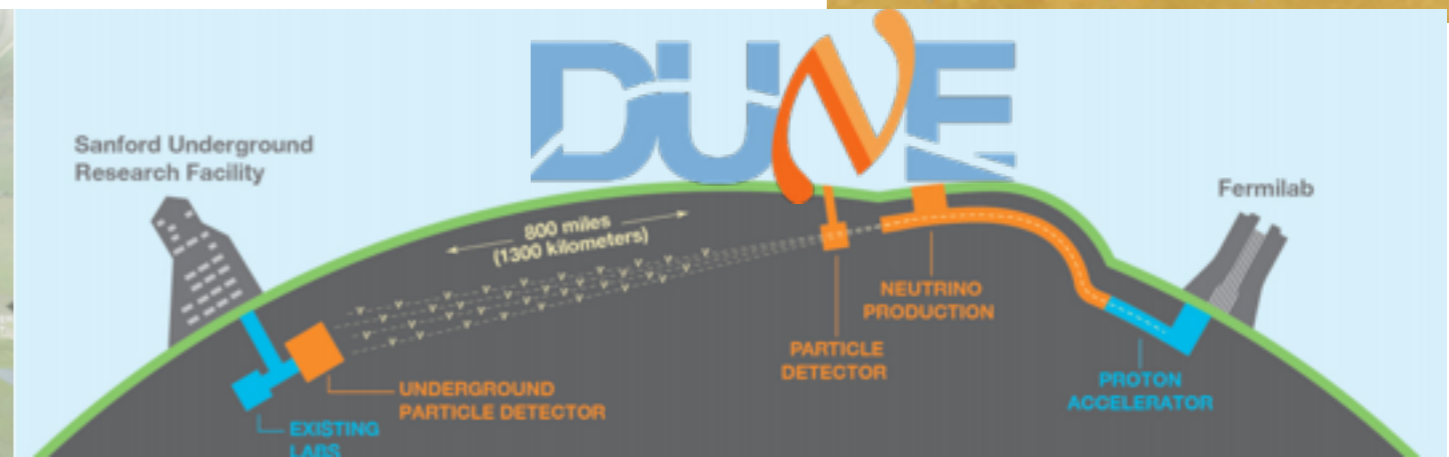
Accelerator based Neutrino Oscillation Experiments

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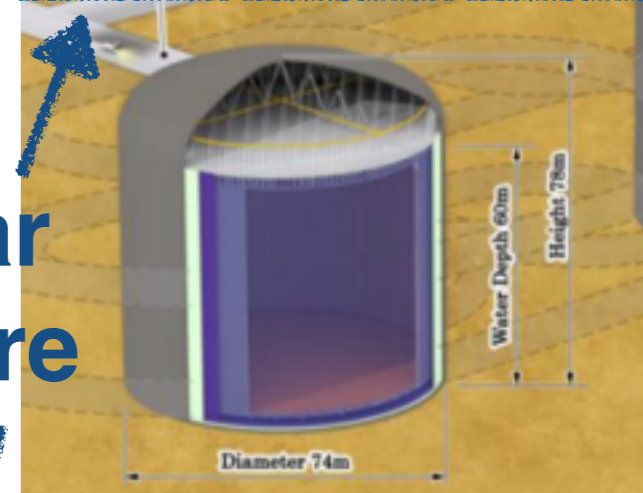
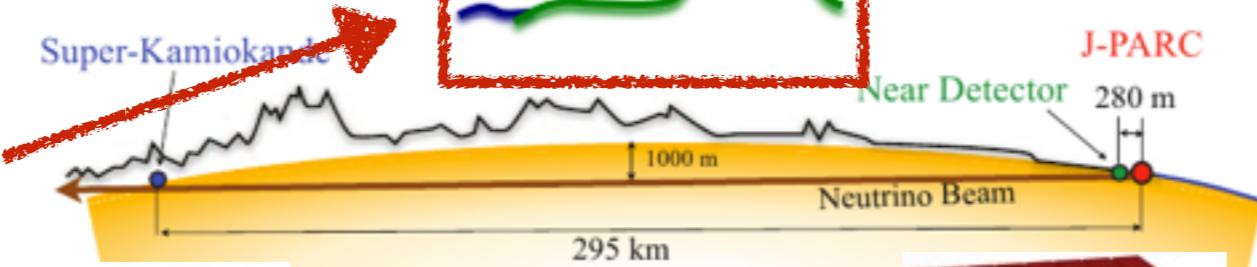
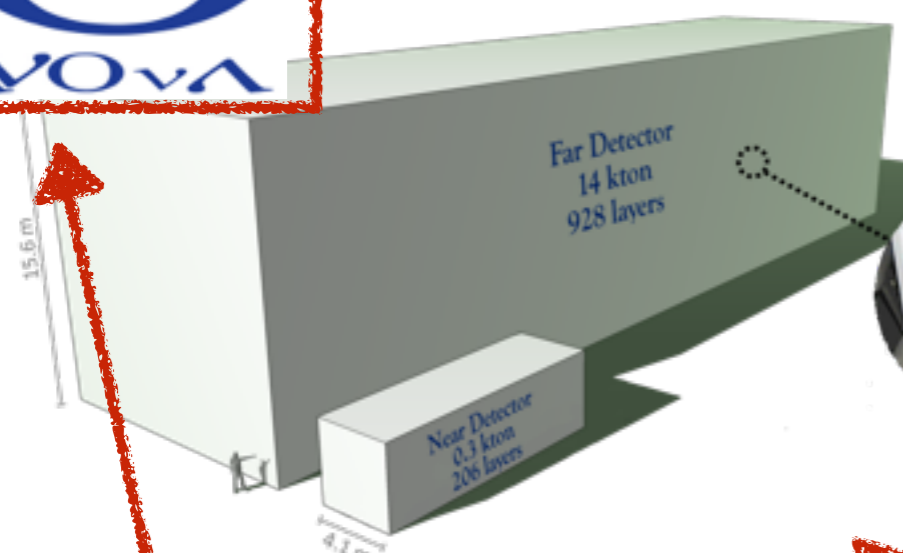
ICARUS

μ BooNE



Accelerator based Neutrino Oscillation Experiments

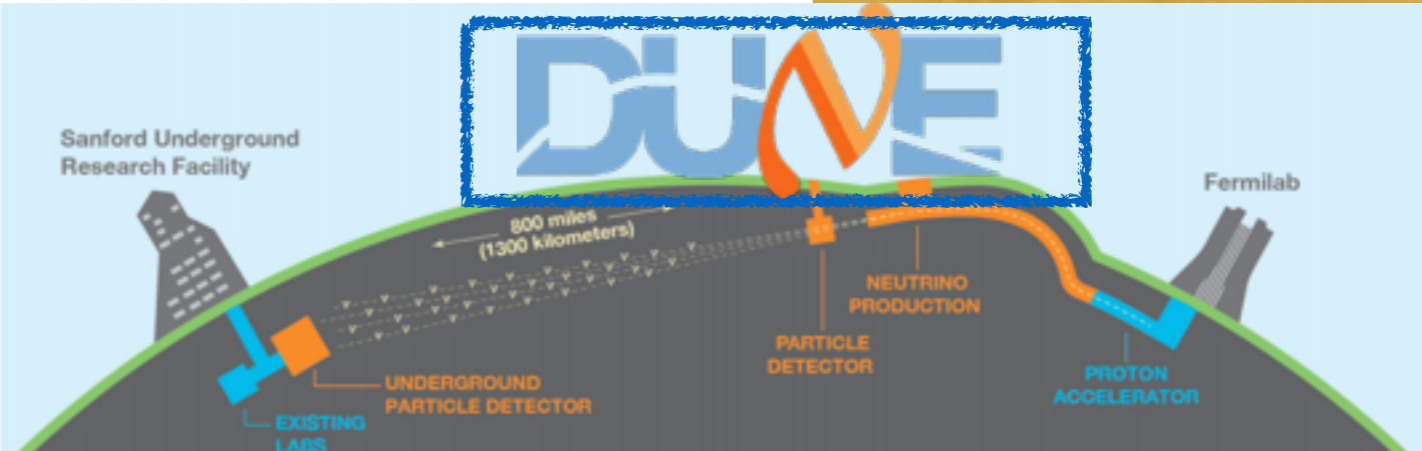
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Near future

Currently running long baseline experiments

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Statistics

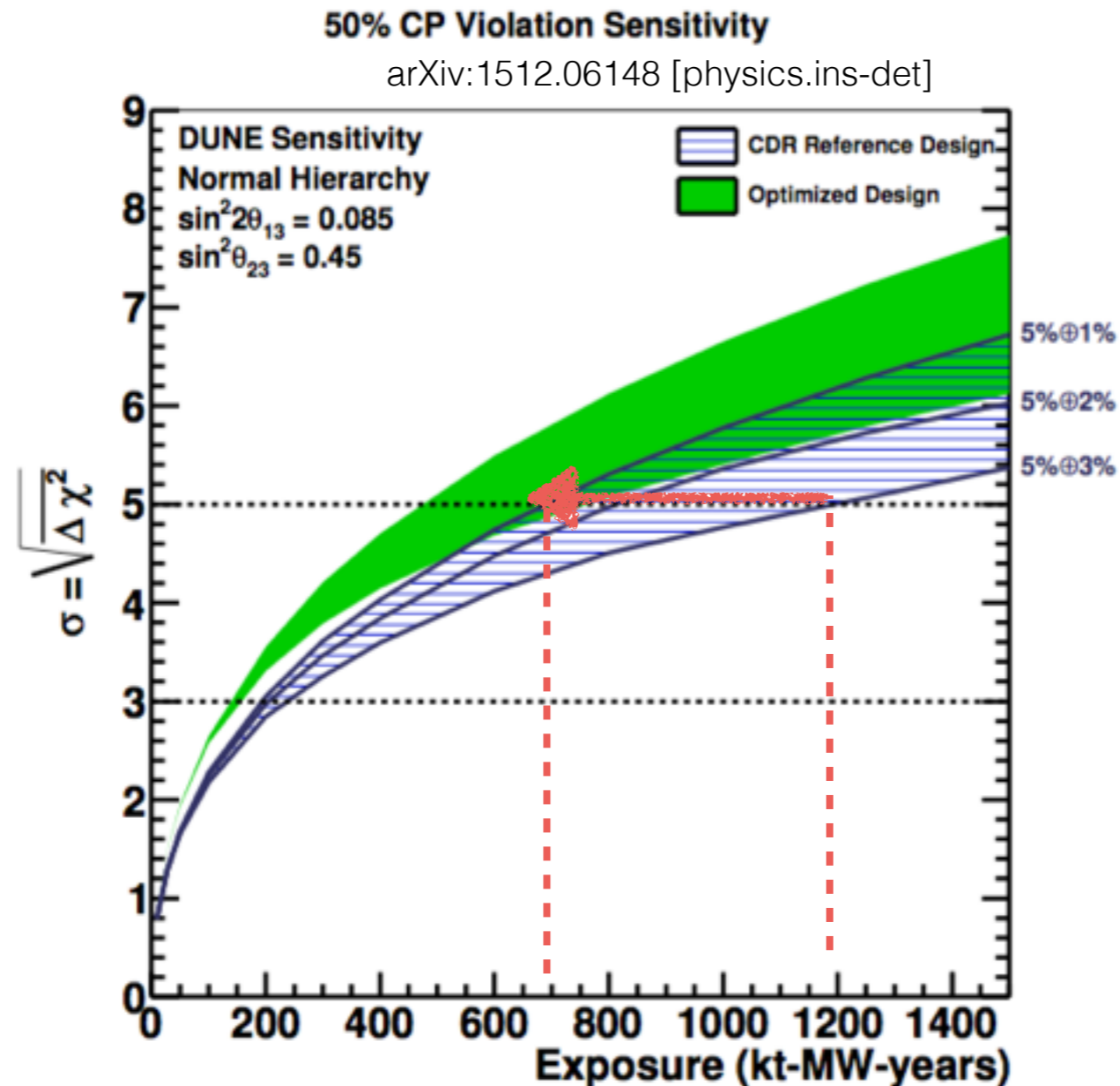
Experiment	$\nu_e + \bar{\nu}_e$	$1/\sqrt{N}$	Ref.
T2K (current)	74 + 7	12% + 40%	2.2×10 ²¹ POT
NOvA (current)	33	17%	FERMILAB-PUB-17-065-ND
NOvA (projected)	110 + 50	10% + 14%	arXiv:1409.7469 [hep-ex]
T2K-I (projected)	150 + 50	8% + 14%	7.8×10 ²¹ POT, arXiv:1409.7469 [hep-ex]
T2K-II	470 + 130	5% + 9%	20×10 ²¹ POT, arXiv1607.08004 [hep-ex]
Hyper-K	2900 + 2700	2% + 2%	10 yrs 2-tank staged KEK Preprint 2016-21
DUNE	1200 + 350	3% + 5%	3.5+3.5 yrs x 40kt @ 1.07 MW arXiv:1512.06148 [physics.ins-det]

Current appearance measurements stats dominate

$O(10^3) \nu_e$ at future experiments → demands ~2% systematics

$O(10^4) \nu_\mu$ → need systematics as good as we can get!

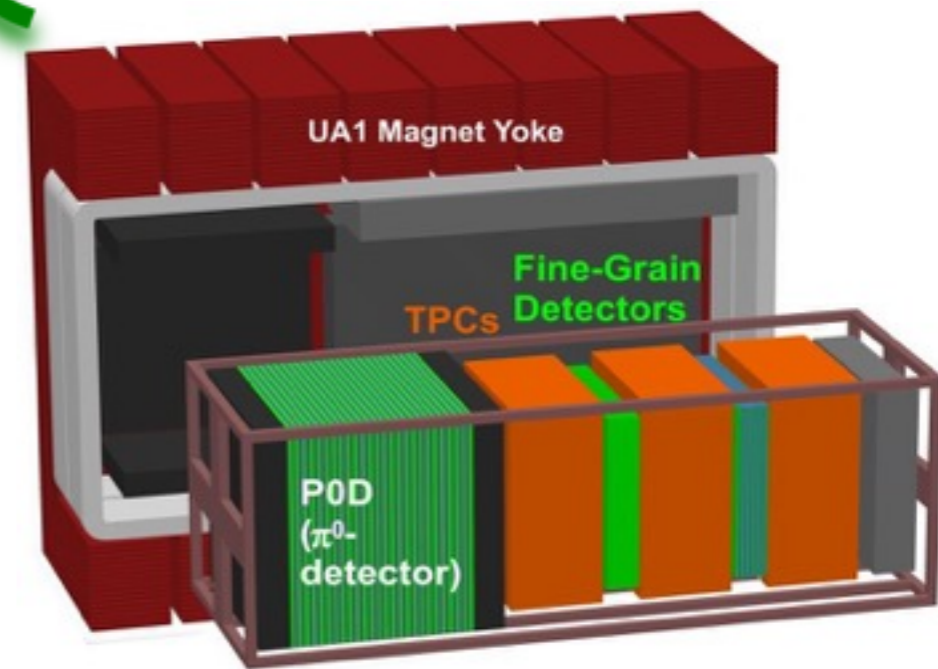
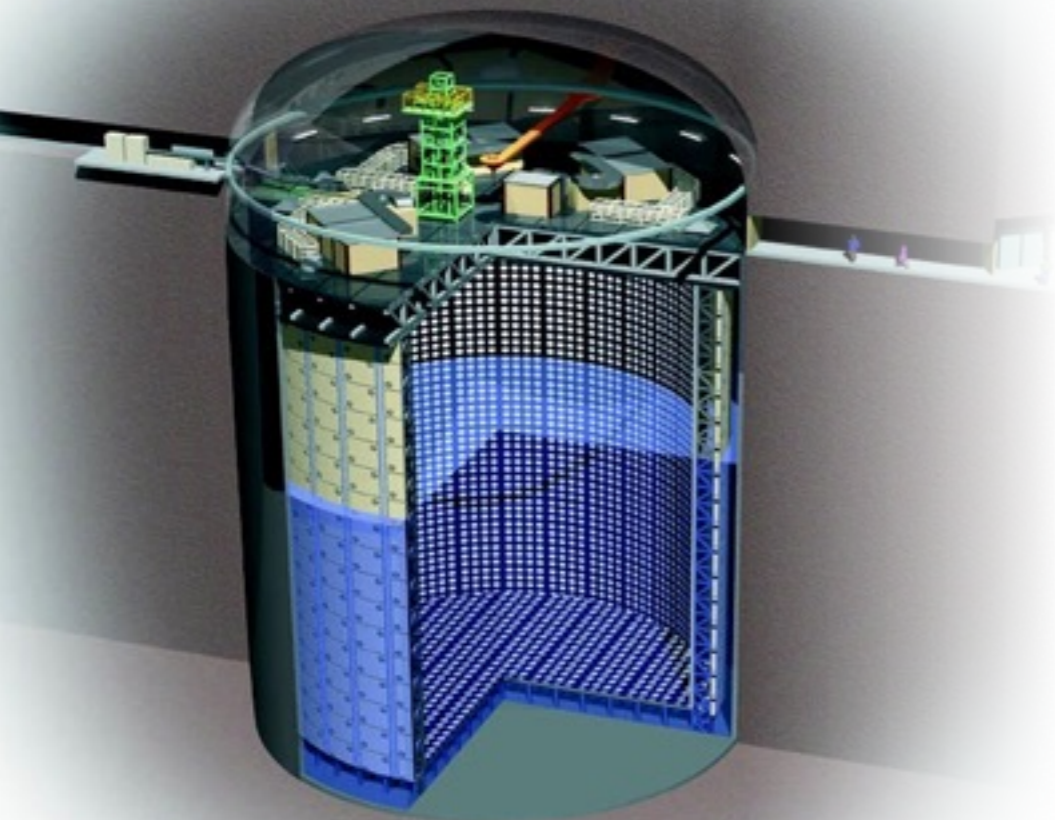
Systematic Uncertainties



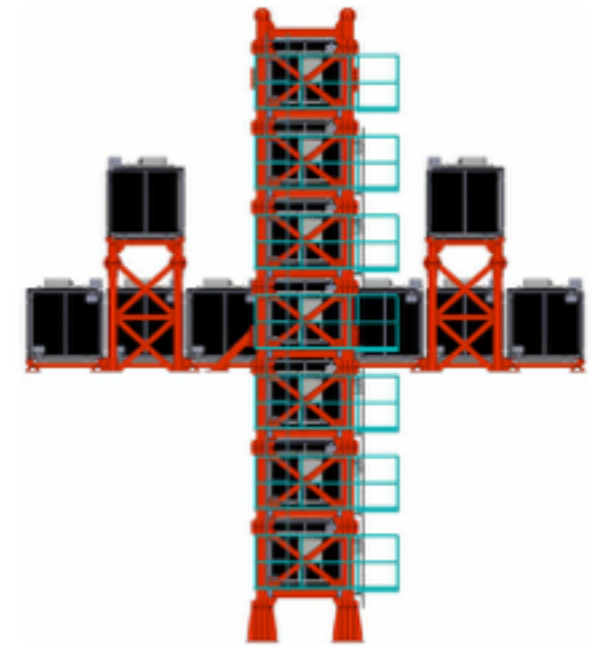
Reduction in systematic uncertainties
can be equivalent to significant boost
in exposure

T2K

 Far Detector
(Super-K)

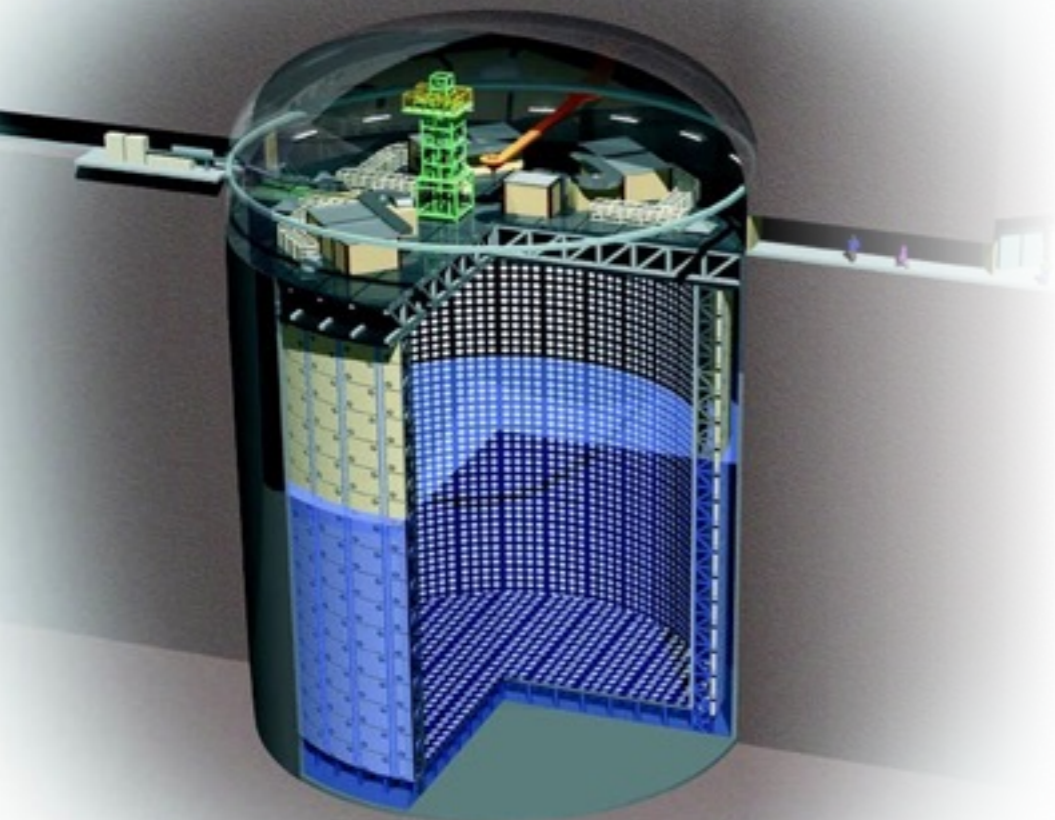


Near Detectors
(ND280+INGRID)



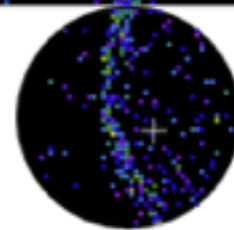
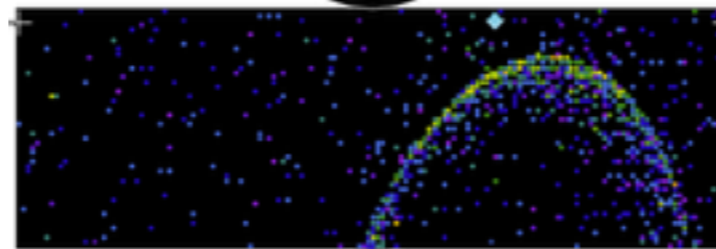
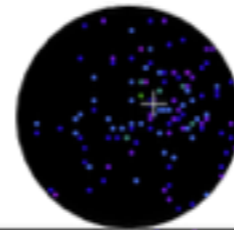


Water Cherenkov Far Detector
>22.5 kt fiducial mass

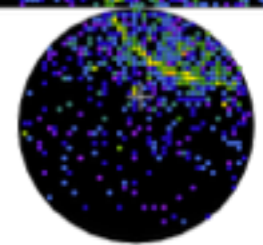
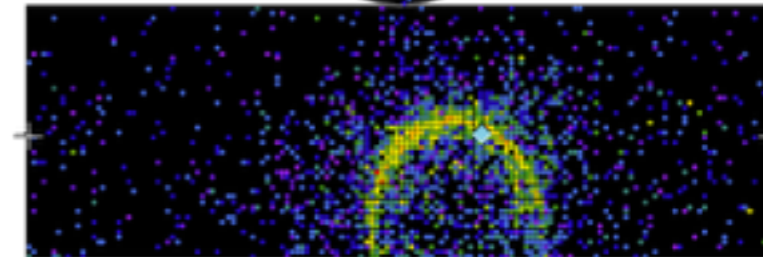
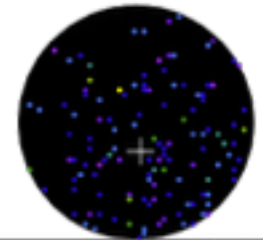


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Muon



Electron



Oxygen target
 4π acceptance

Energy reconstruction from lepton kinematics

Blind to particles below Cherenkov threshold
for protons < 1.1 GeV/c.

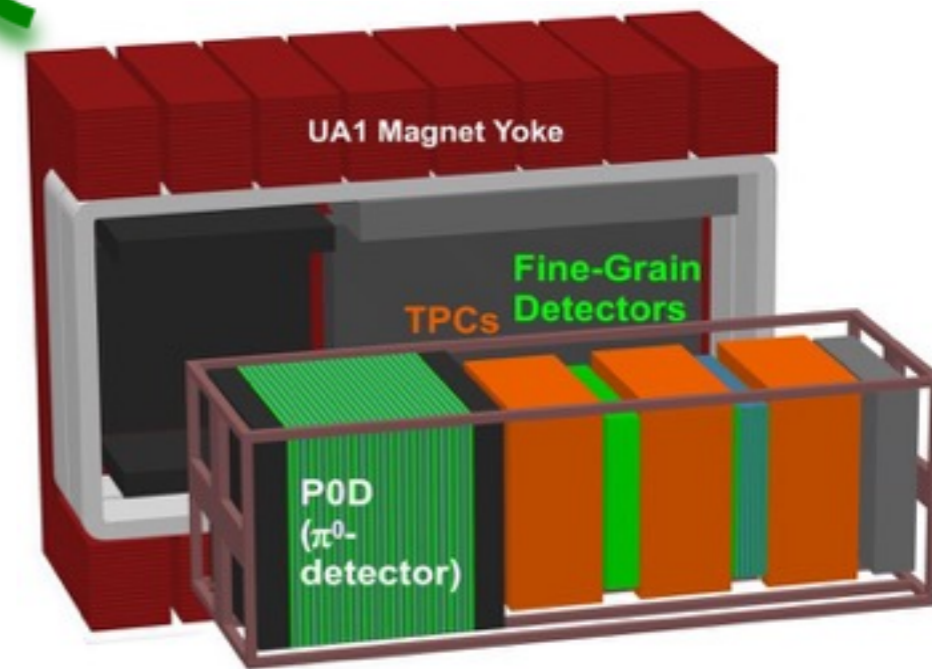
(neutron counting possible with SK-Gd)

T2K

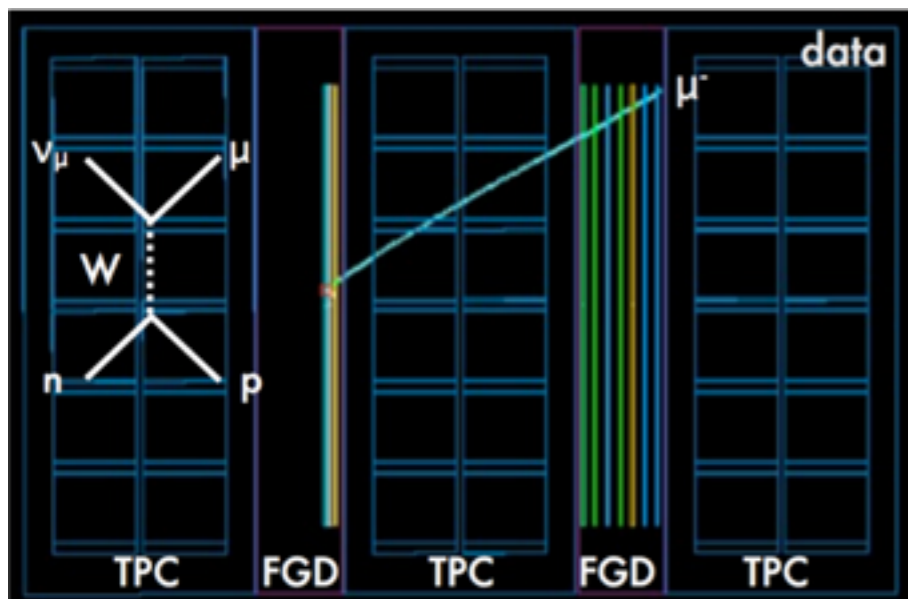
Carbon and Oxygen target materials

Acceptance differs from far detector

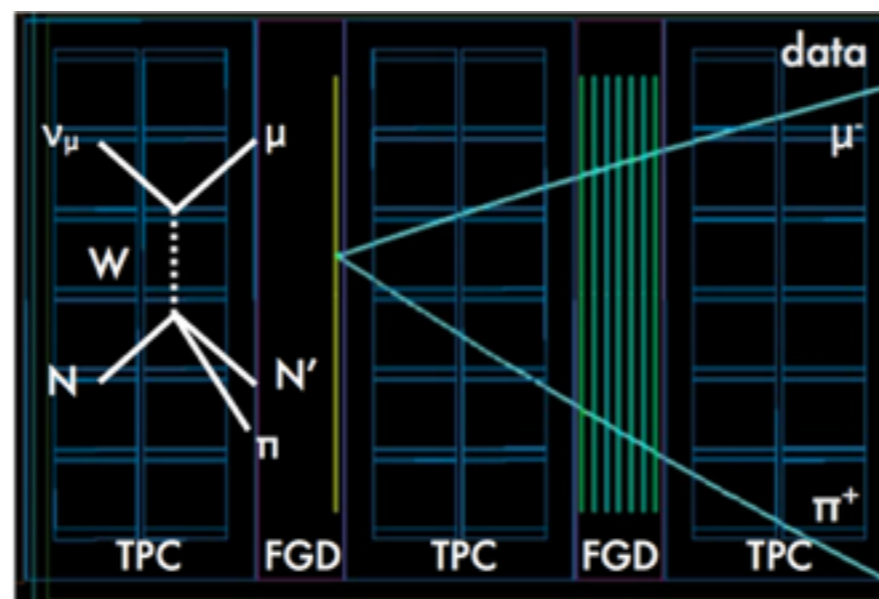
Magnetic field for sign selection



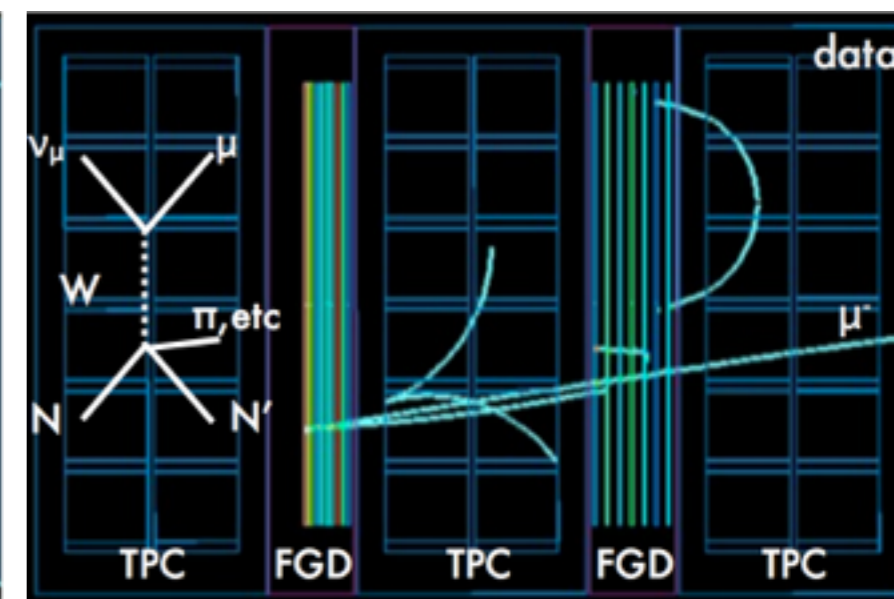
Near Detector (ND280)



CC $1\mu + 0\pi + X$

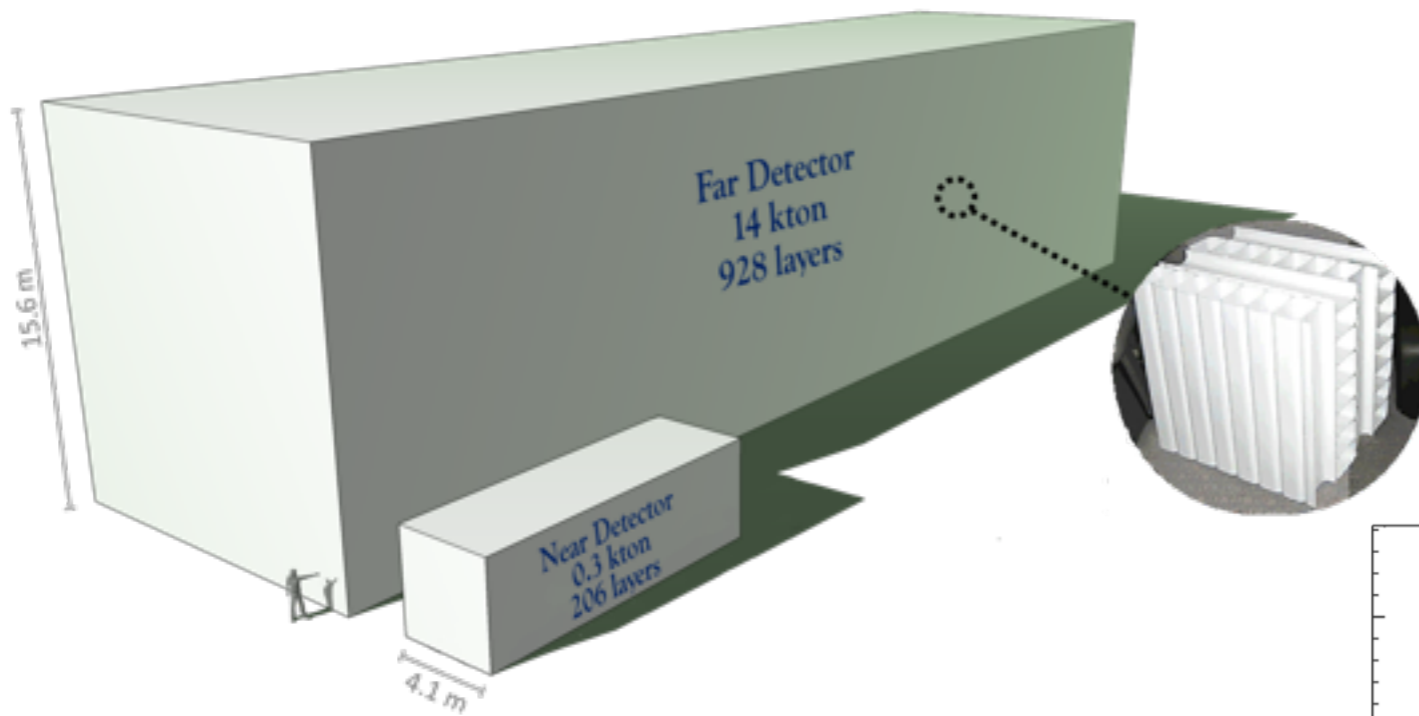


CC $1\mu + 1\pi^+ + X$



CC other

NOvA Experiment

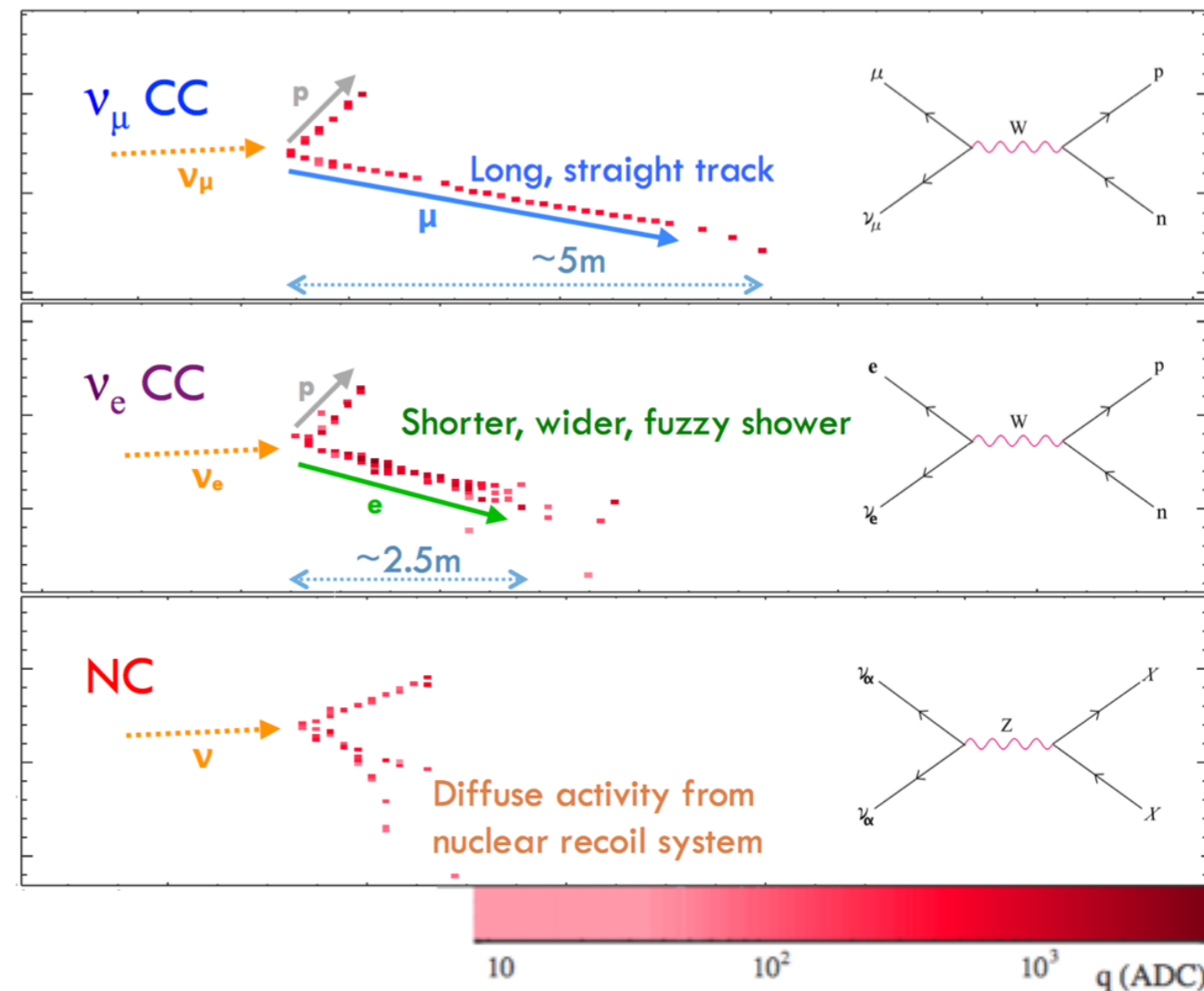


Liquid Scintillator tracking calorimeter

Almost identical near far detectors

Calorimetric Energy Reconstruction

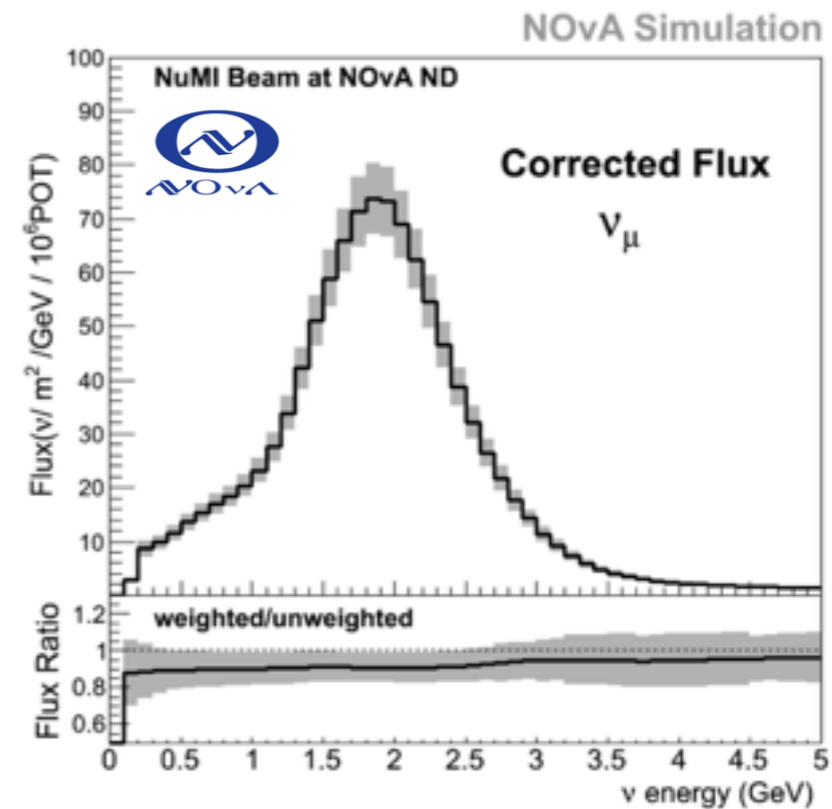
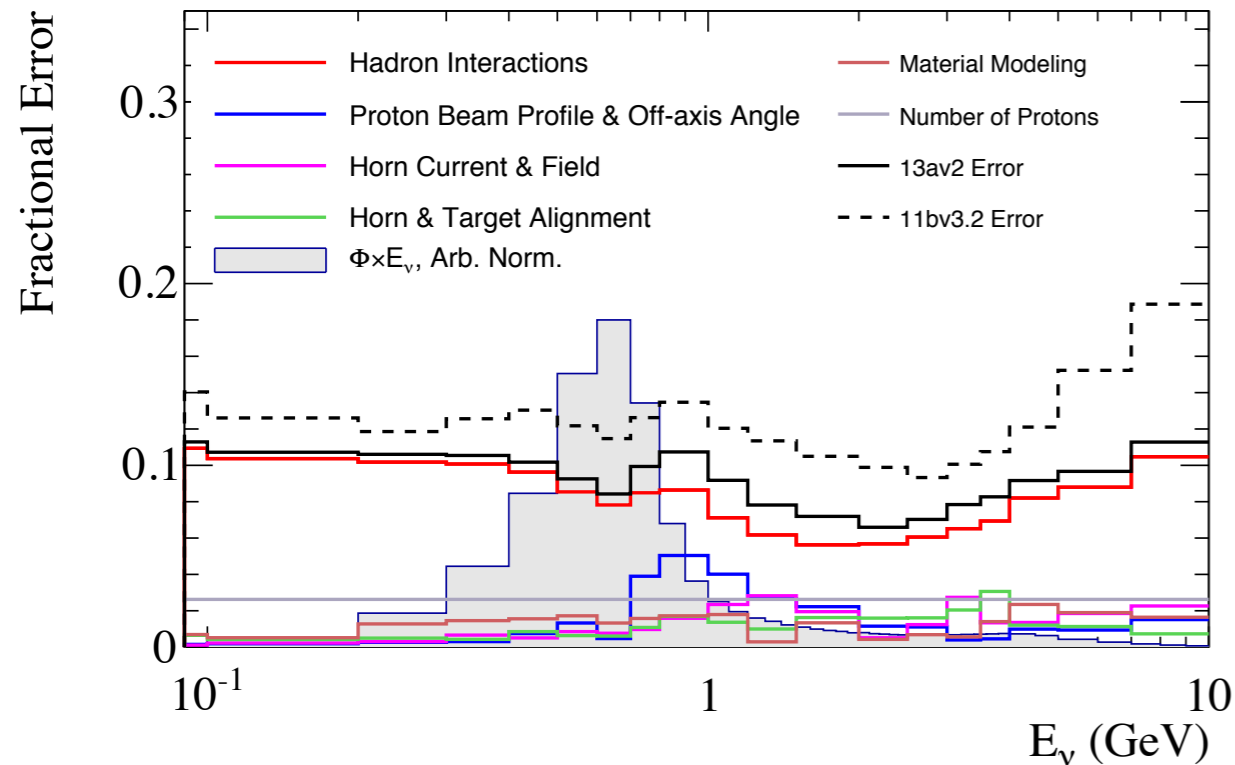
$$E_{\nu \text{ reco}} = E_{\text{lep}} + E_{\text{had}}$$



Flux Uncertainties

SK: Neutrino Mode, ν_μ

T2K



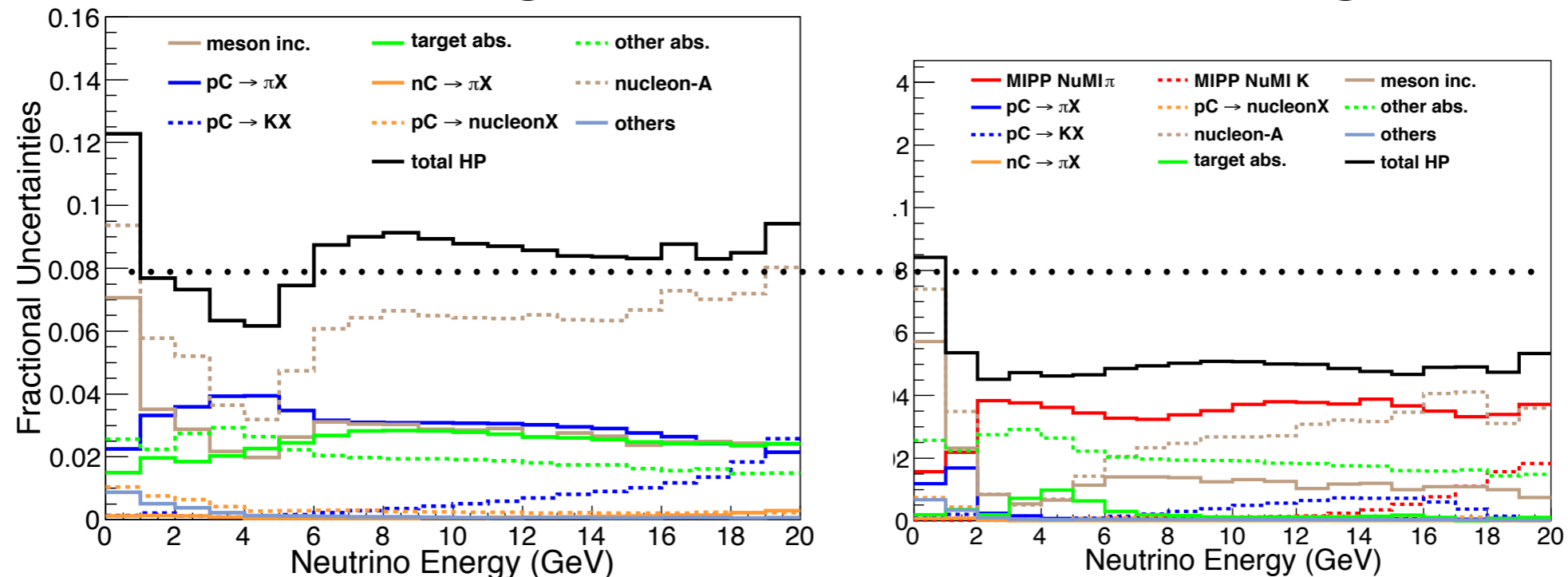
T2K \sim 8-12% (based on thin target tuning)

Dominated by hadron interaction modelling

Alignment/focussing uncertainties are also important
(especially for near to far extrapolation)

Flux Uncertainties

Thin Target \longrightarrow Thick Target



MINERvA Low E NuMI Flux Uncertainties, Phys. Rev. D 95, 039903 (2017)

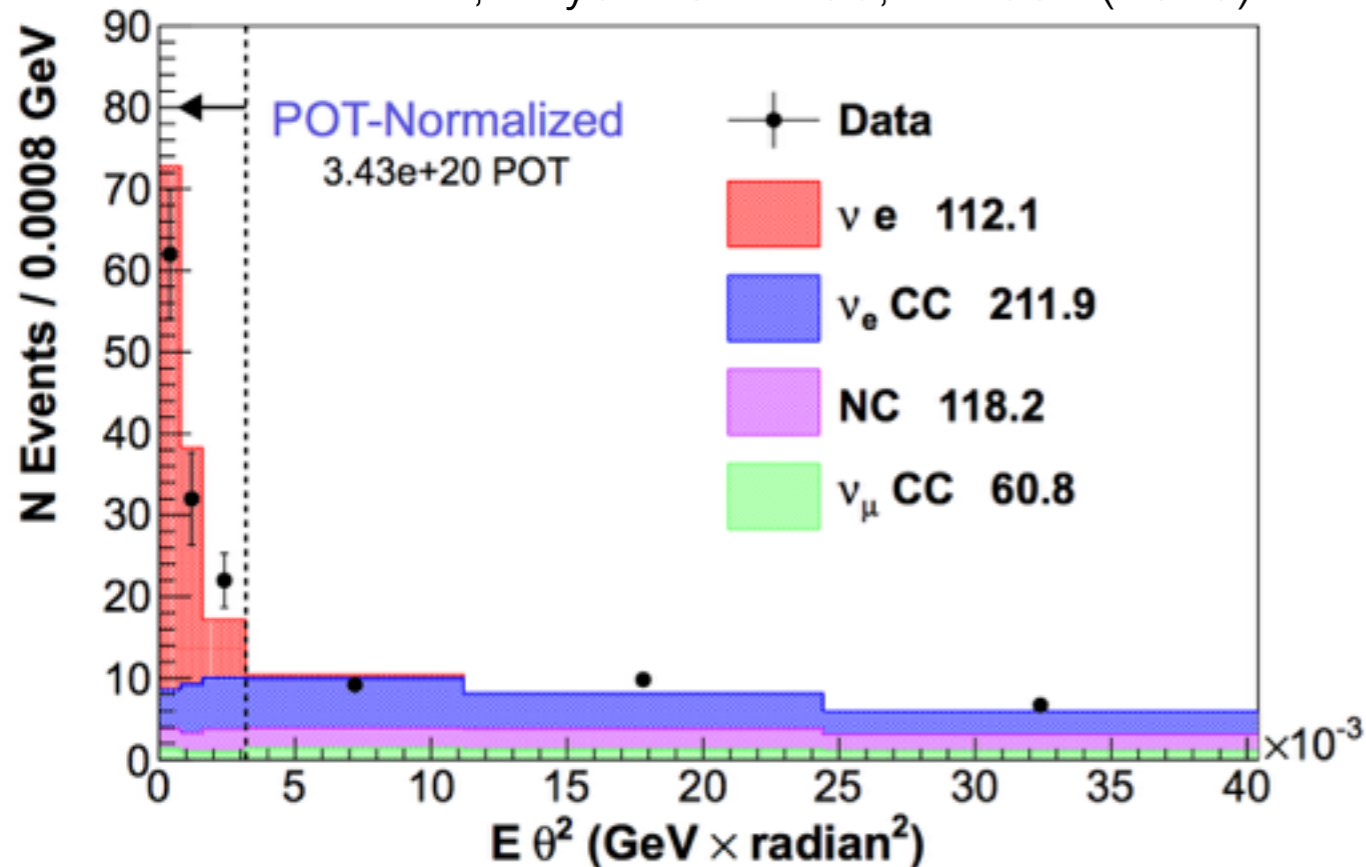
Significant reductions from thick/replica target
(See Tomislav Vladislavljevic poster for latest T2K tuning)

Future high beam power experiments may have different target material/geometry requiring dedicated hadron production measurements

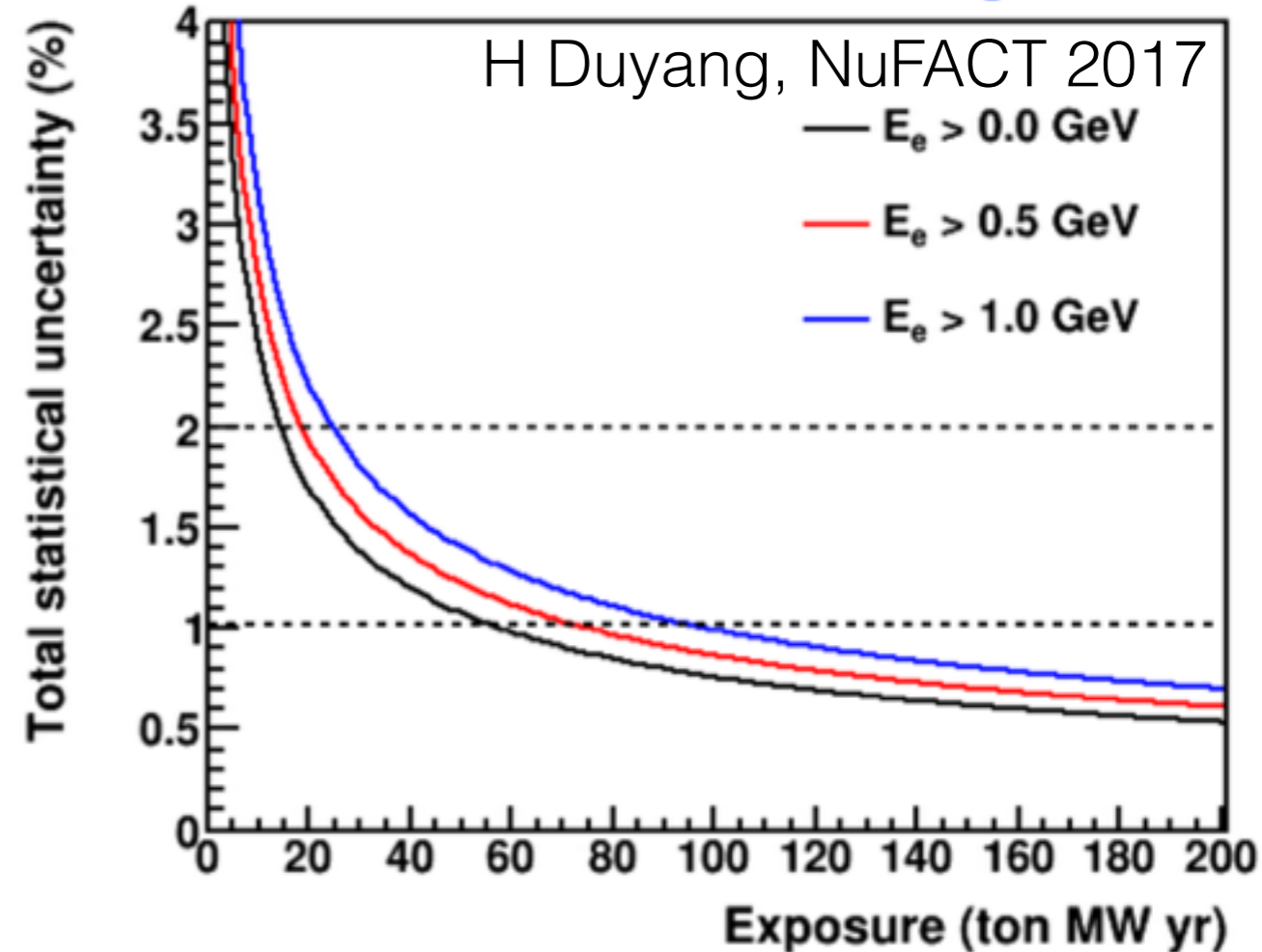
Flux Uncertainties

In situ flux measurements possible
eg neutrino-electron elastic scattering

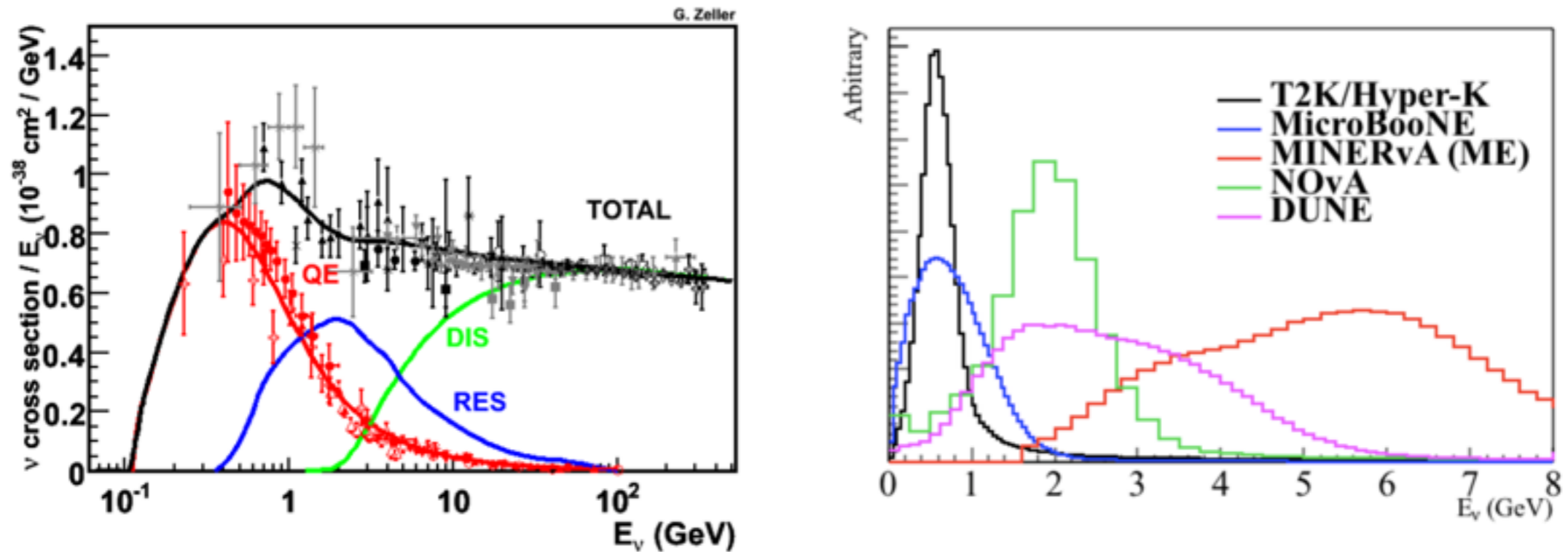
MINERvA, Phys. Rev. D 93, 112007 (2016)



DUNE Work in Progress



Neutrino Interaction Model Uncertainties



Wide range of processes need to be simulated
Require both lepton and hadronic side of the interaction
Nuclear effects important in the relevant energy regime

Experiments rely on MC generators
for $E_{\text{visible}} \rightarrow E_{\nu}$ extrapolation

Model parameter uncertainties from fits to external datasets

Sometimes parameter error must be inflated or ad-hoc parameters to account
for discrepancies between model and data or known flaws in the model

T2K Cross-Section Model

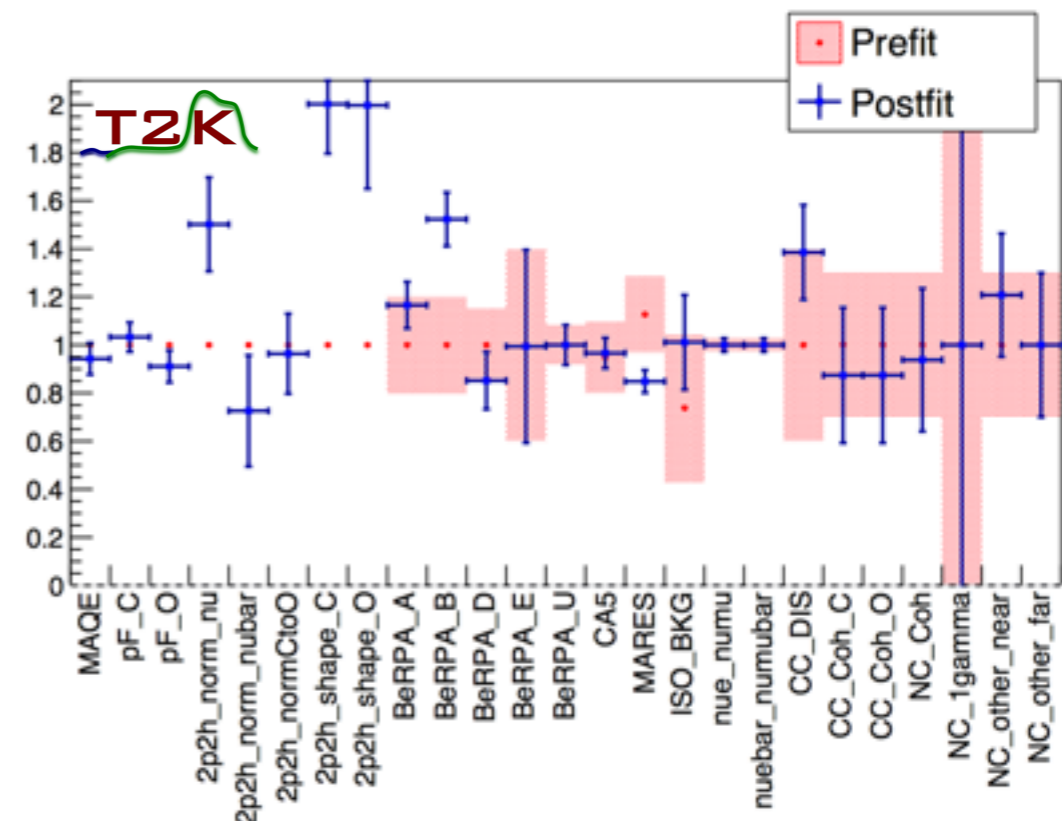
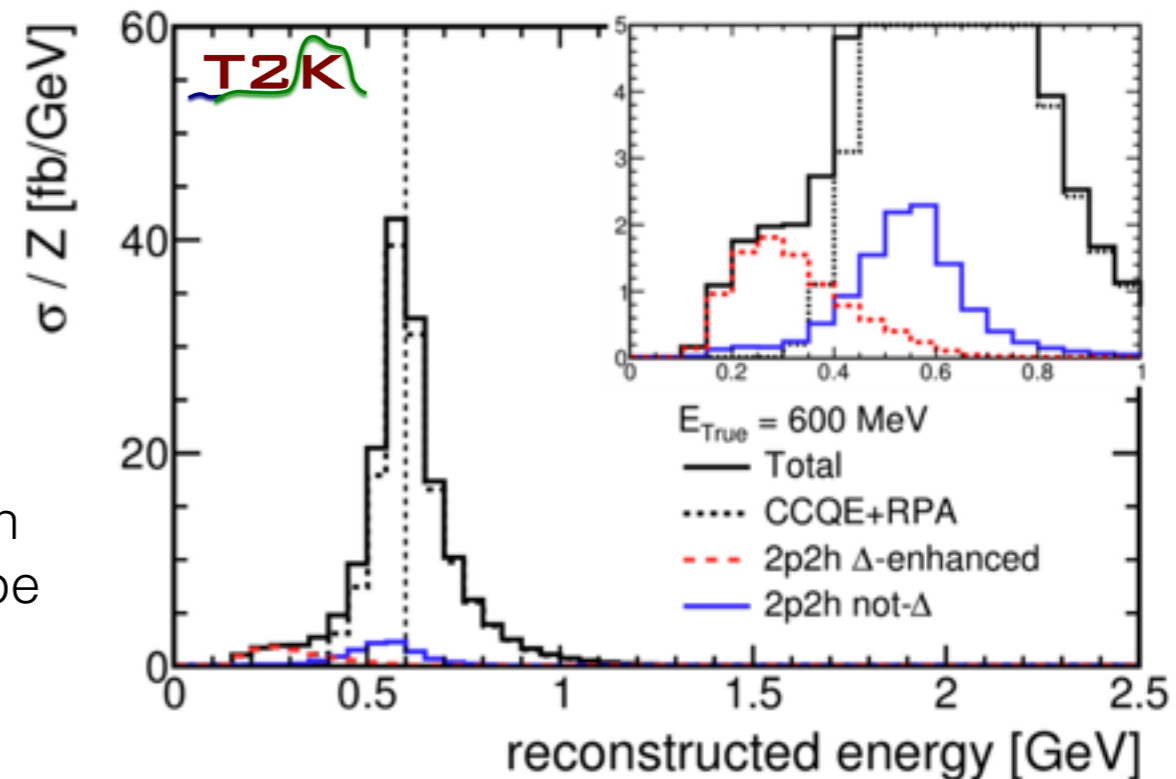
Implemented in NEUT MC generator

Quasi-elastic scattering most important process at T2K energies

- Valencia 2p-2h model Phys. Rev. C83 (2011) 045501
- Long-range effects with Random Phase Approximation
- Parameters introduced to vary normalisation and shape
- Relativistic Fermi Gas (RFG) nuclear model
- Uncertainties from RFG \leftrightarrow Local Fermi Gas
- Final state interactions with cascade model

No priors on most CCQE parameters
Constraint from near detector

Impact of alternative models not implemented in oscillation analysis evaluated with fake data studies

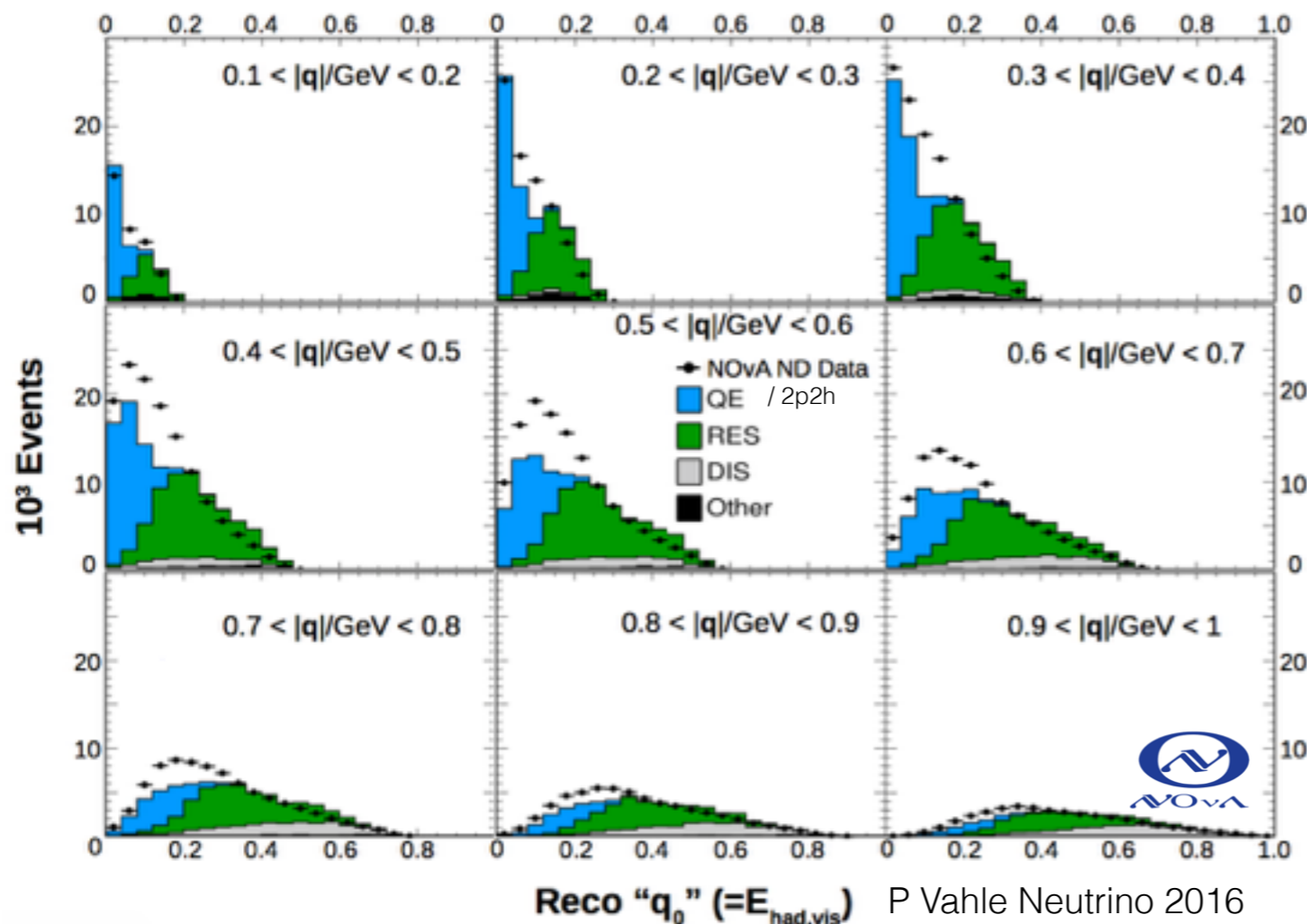


NOvA Cross-section Model

Use GENIE MC generator and uncertainties

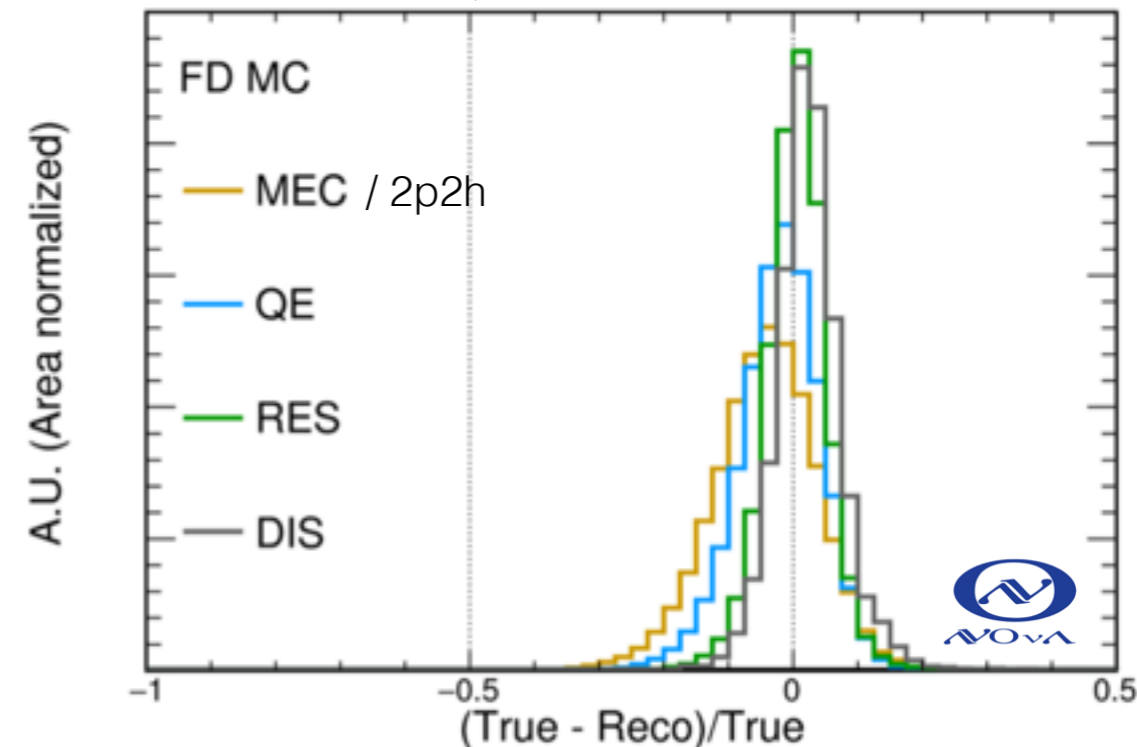
Some additions/modifications

- Empirical 2p2h model, tuned to match ND data
- Parameters to cover RPA uncertainties
- Alternative tuning of CC1 π model [Eur. Phys. J. C 76, 474 (2016)]

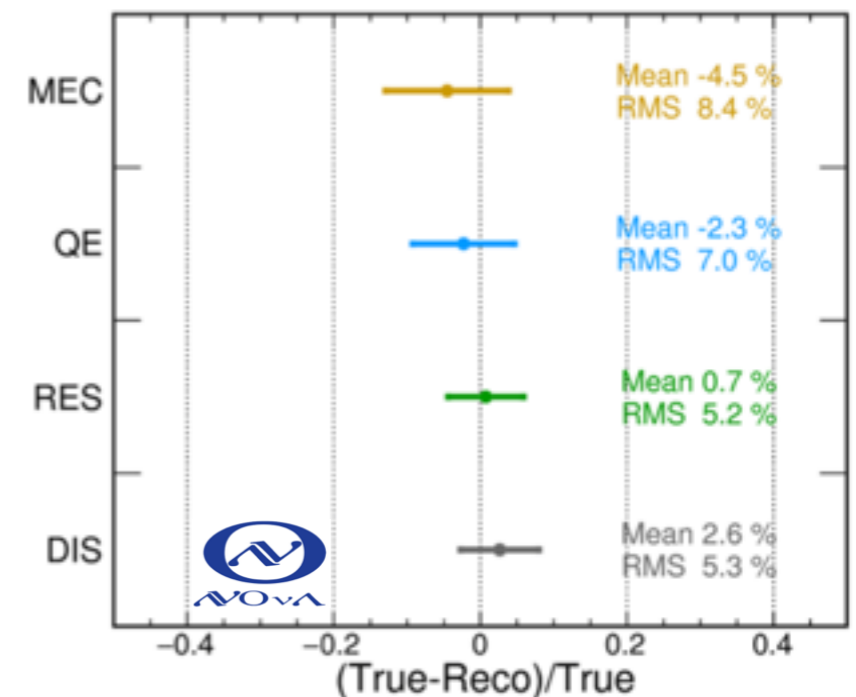


P Vahle Neutrino 2016

J. Wolcot, NuInt 2017 NOvA Simulation



NOvA Simulation

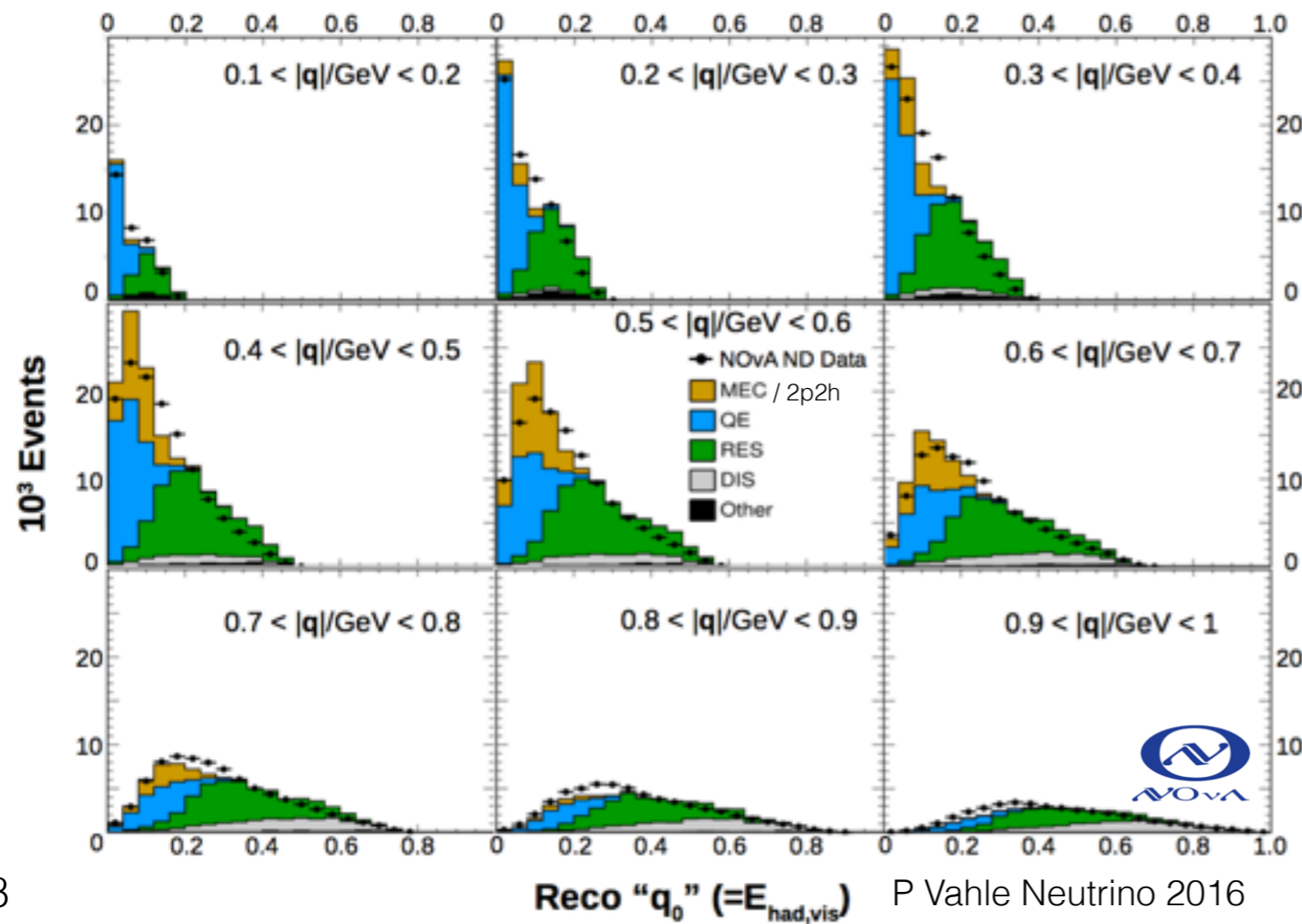


NOvA Cross-section Model

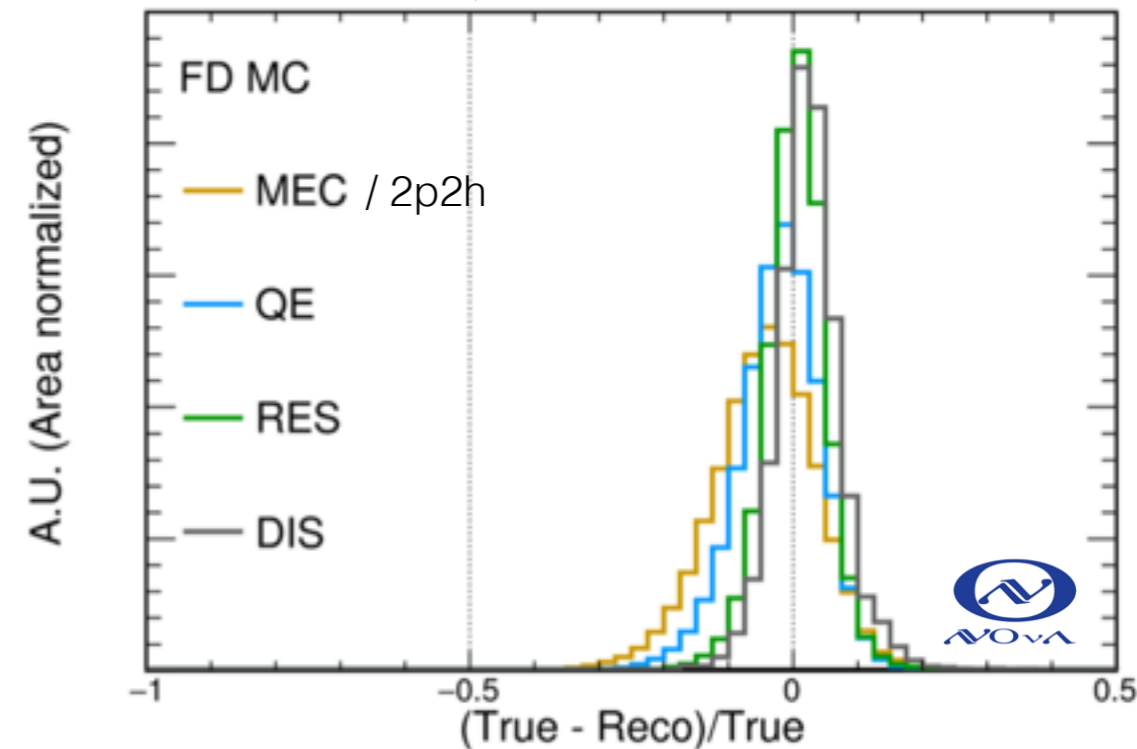
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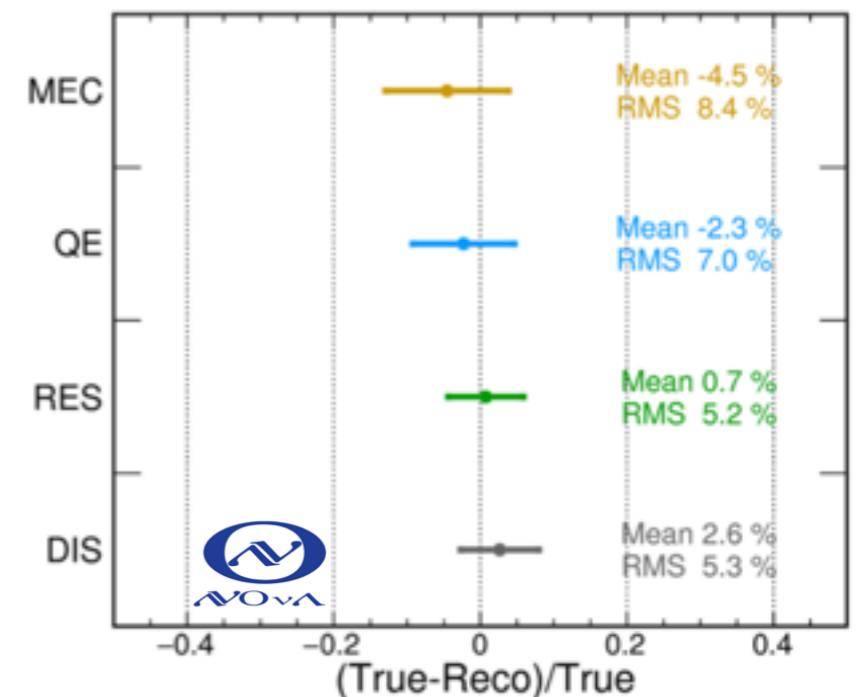
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J. Wolcot, NuInt 2017 NOvA Simulation

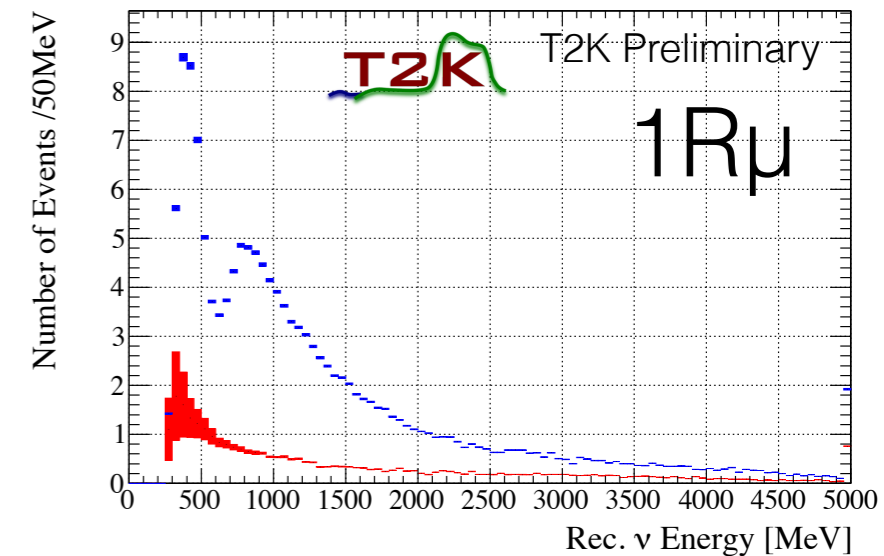
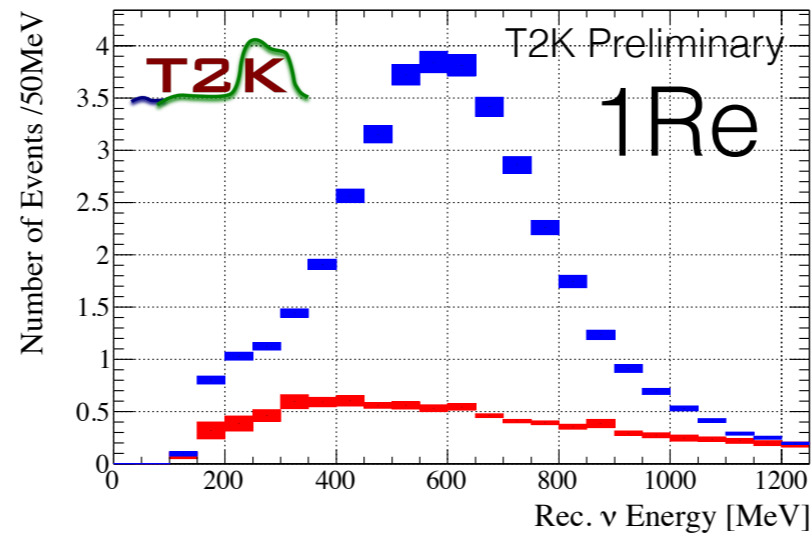


NOvA Simulation

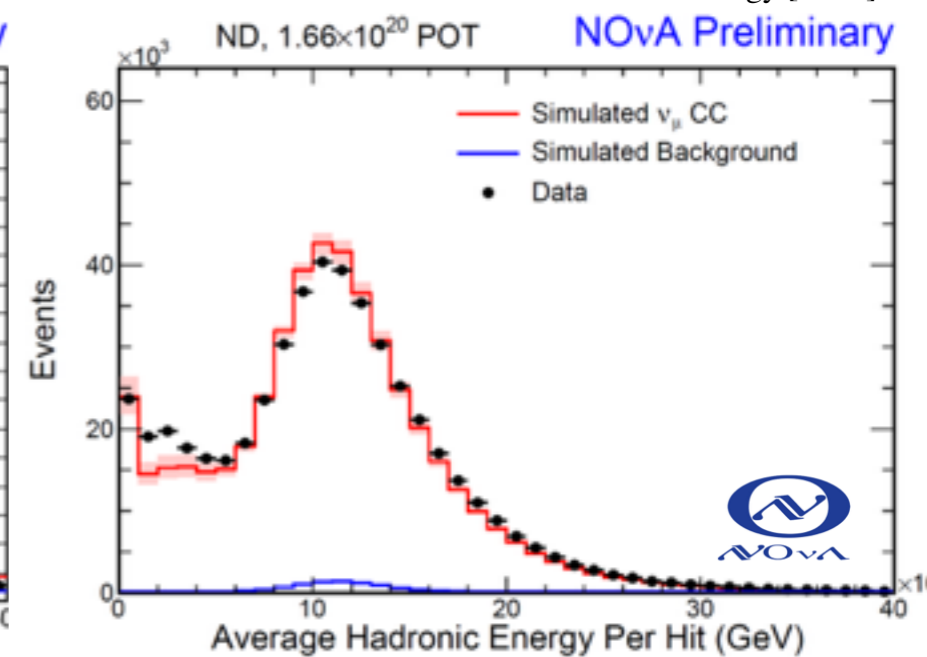
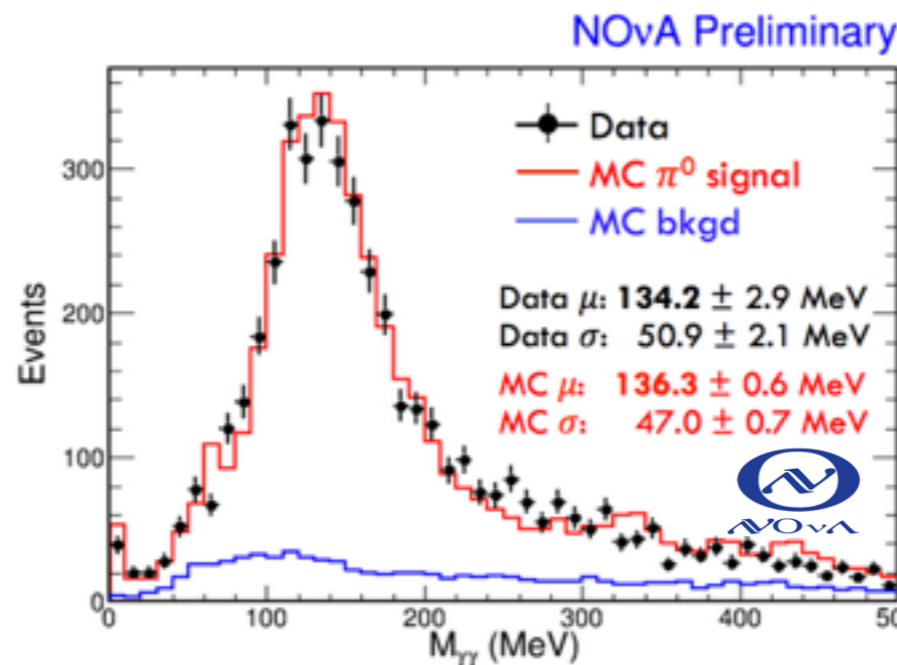


Detector Modelling Uncertainties

SK detector response evaluated with atmospheric sample



NOvA detector response evaluated with beam and cosmic muon samples in both ND and FD



Detector modelling uncertainties typically from data MC comparisons in control samples
 May be limited by control sample statistics

T2K Systematic Uncertainties

ND280 constraint
13% → 3%

Error Source	μ sample [%]		e sample [%]	
	ν	$\bar{\nu}$	ν	$\bar{\nu}$
SK Detector	1.9	1.6	3.0	4.2
SK FSI+SI+PN	2.2	2.0	2.9	2.5
ND280 Constraint (Flux + Cross Section)	3.3	2.7	3.2	2.9
$\sigma(\nu_e)/\sigma(\nu_\mu)$	-	-	2.6	1.5
NC 1γ	-	-	1.1	2.6
NC other	0.3	0.1	0.1	0.3
Total Systematic	4.4	3.8	6.3	6.4
Statistical	6.5	12	12	40

T2K preliminary (final systematics pending)

Total systematic uncertainty
~4 - 6%

Smaller than stats. uncertainty
(for now!)

Pion Final State
Interactions (FSI) and
Secondary Interactions
(SI) modelling important

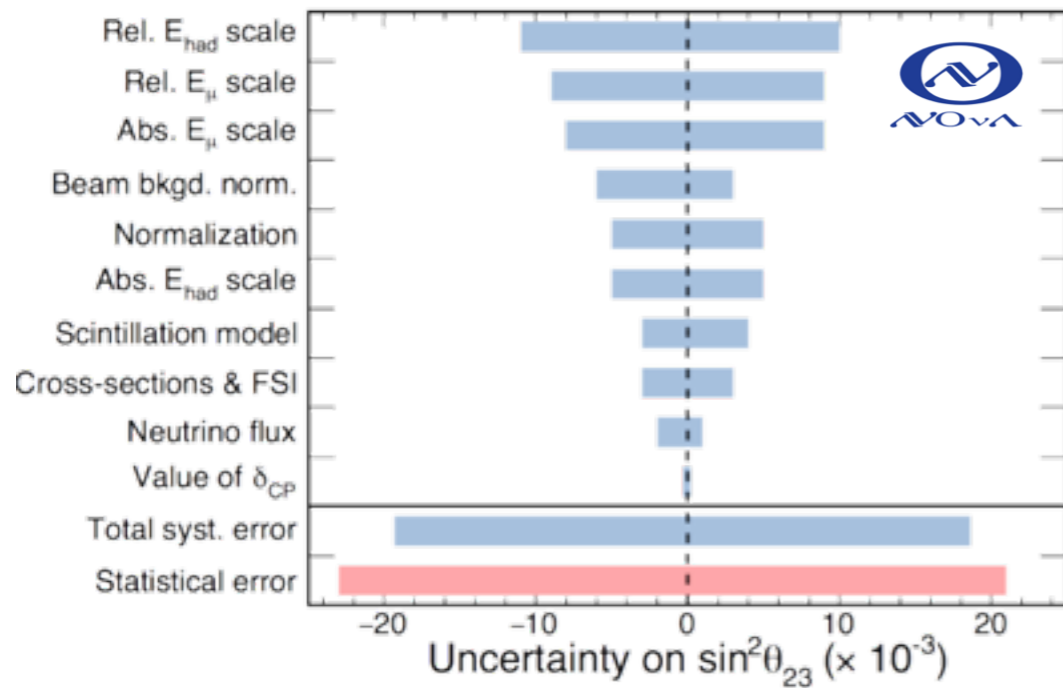
Theoretical uncertainty

ν_e to ν_μ

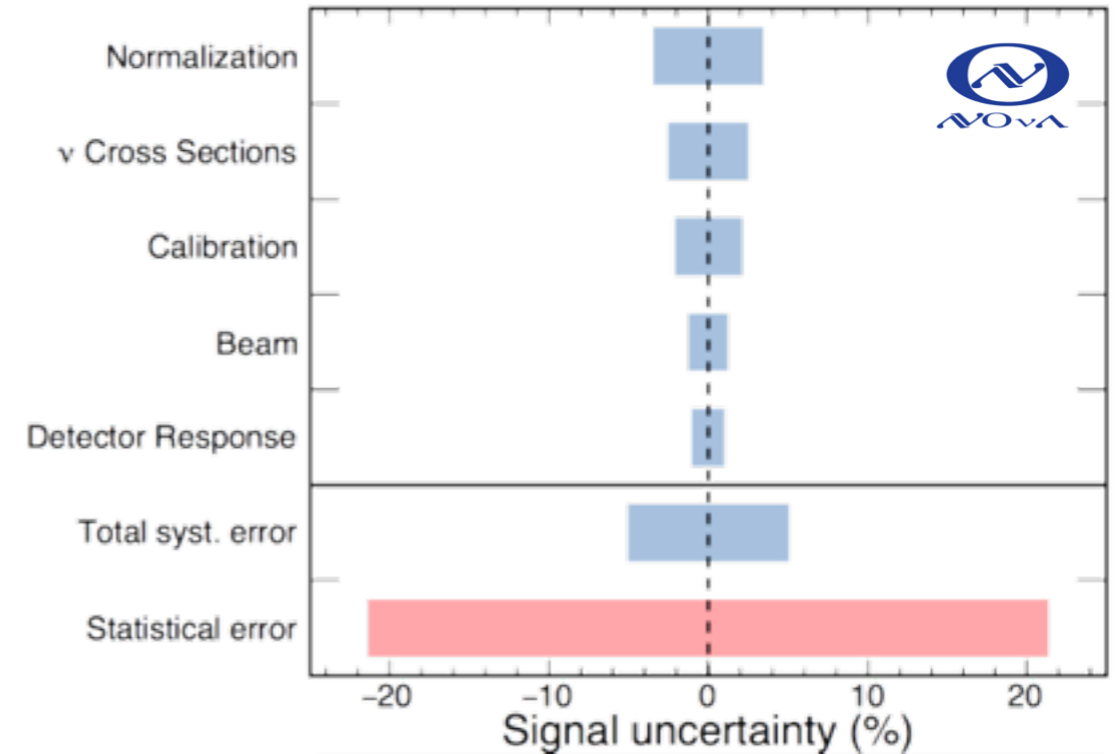
Difficult to constrain with
near detector

NOvA Systematic Uncertainties

ν_μ



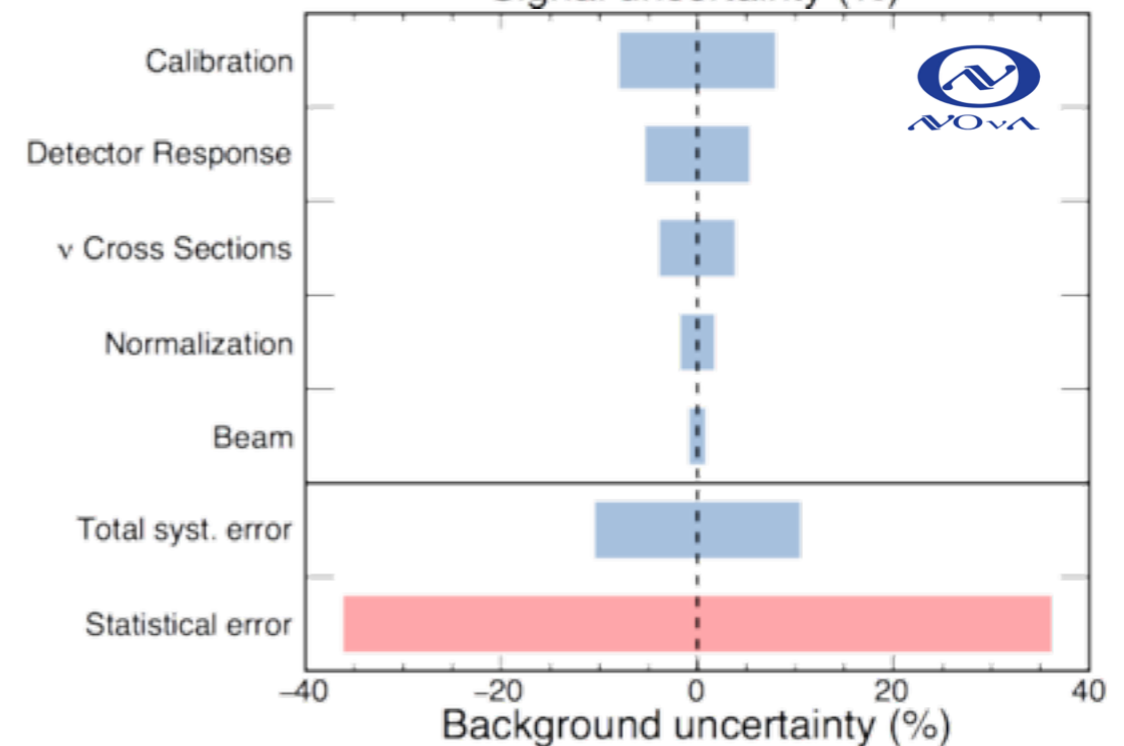
ν_e



$\nu_e \sim 5-10\%$

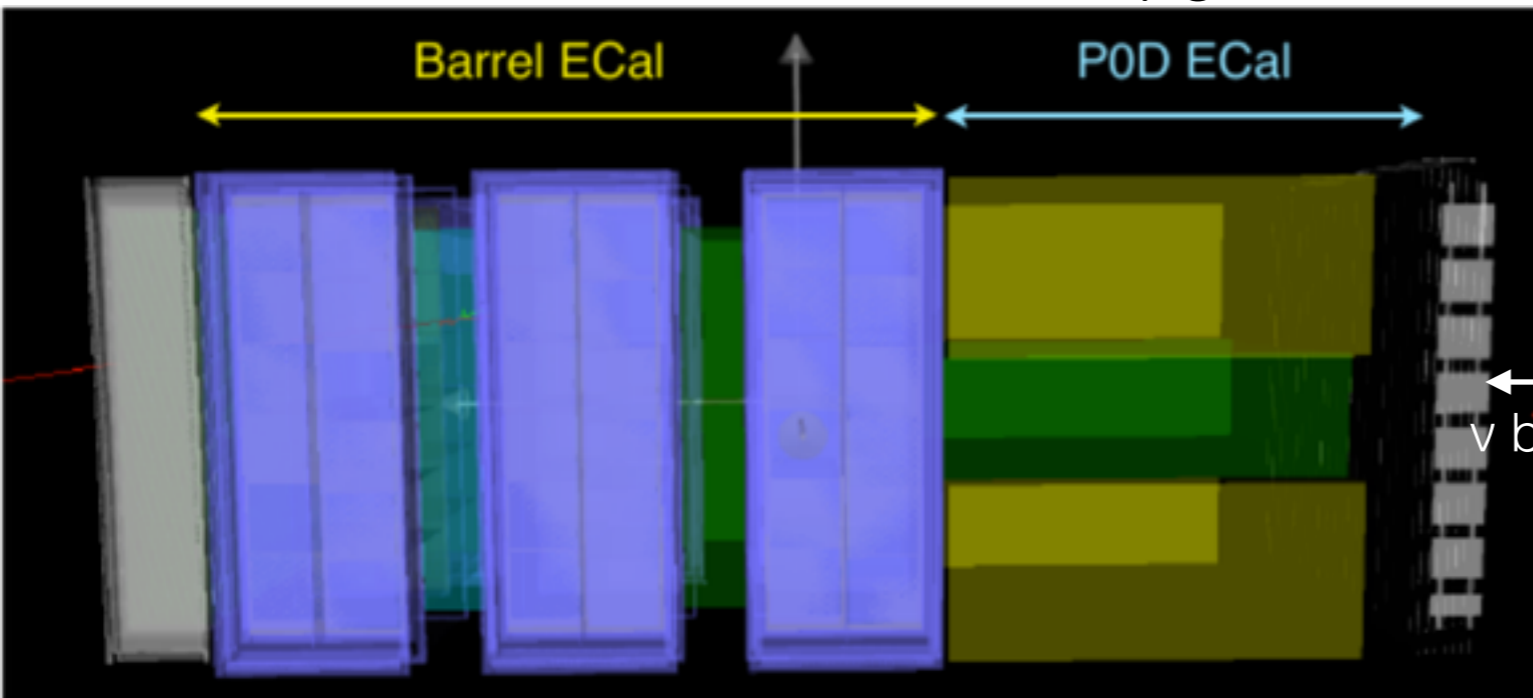
$\nu_\mu \sim 3-4\%$

Energy scale uncertainties dominate ν_μ

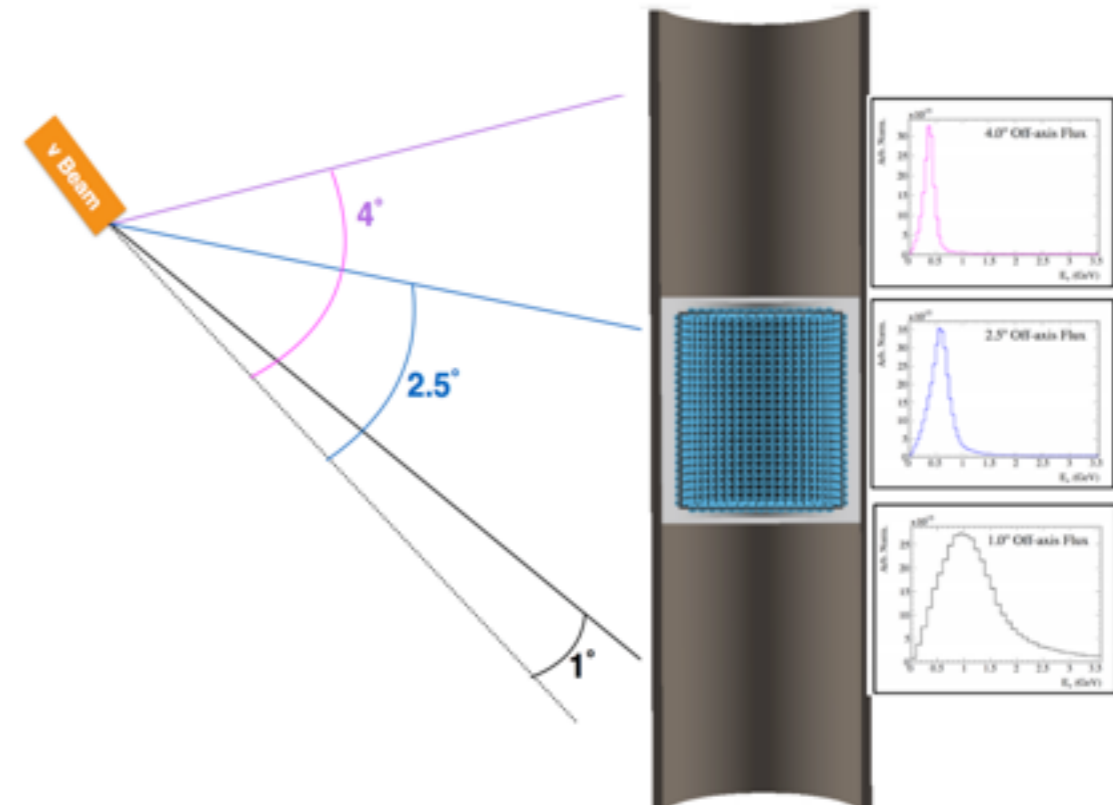


Near Detector Development

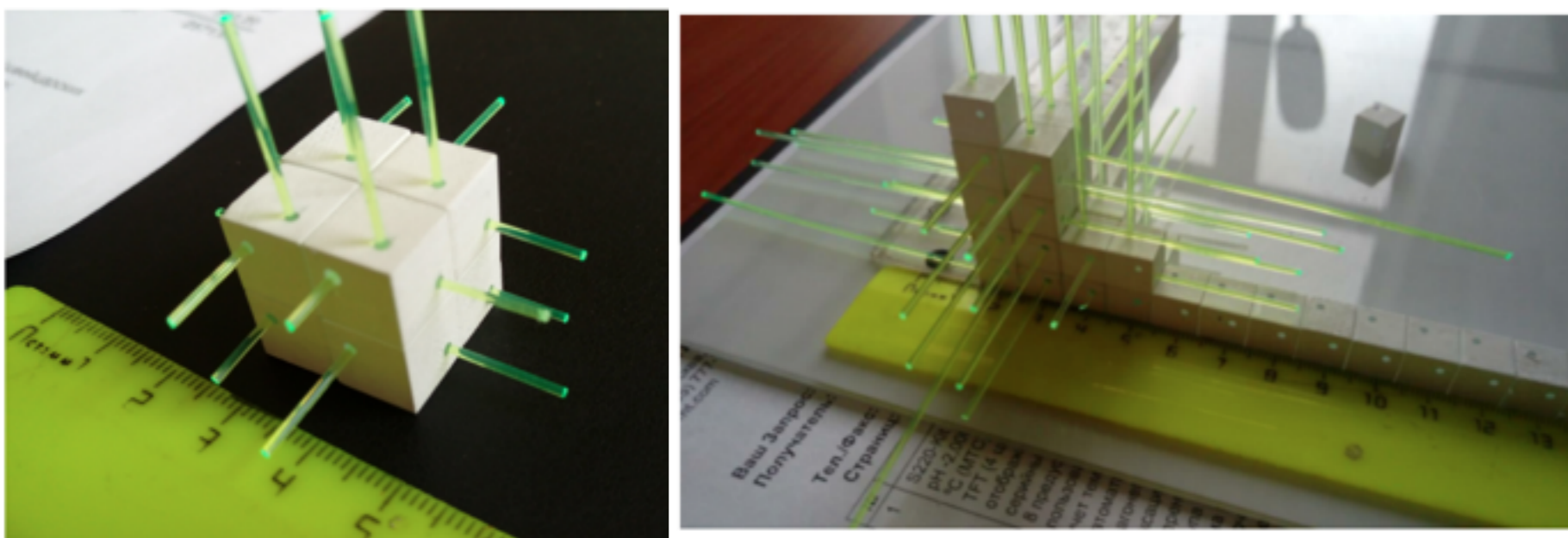
Planned ND280 Near Detector Upgrade



E61 Experiment



Near detector upgrades for T2K-II and T2HK era
New target with increased angular acceptance

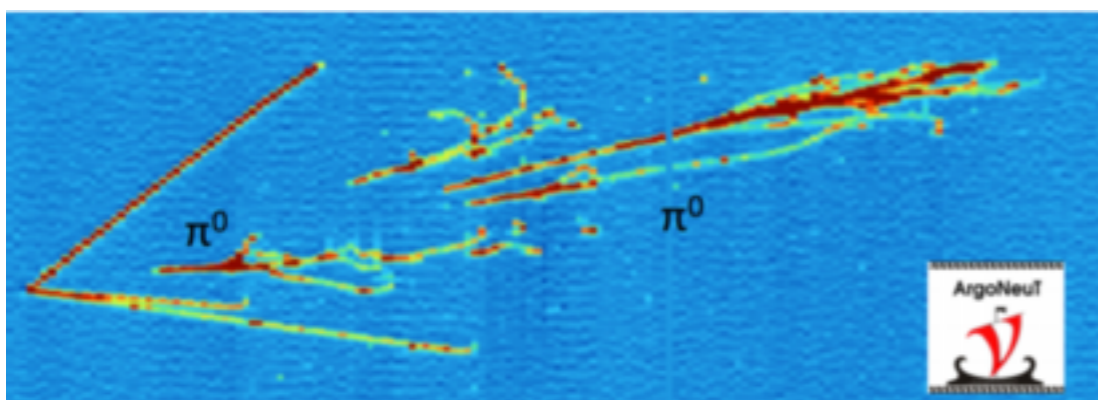
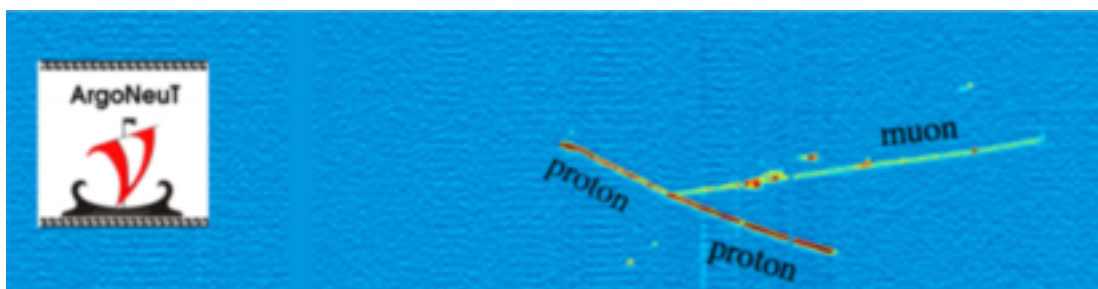
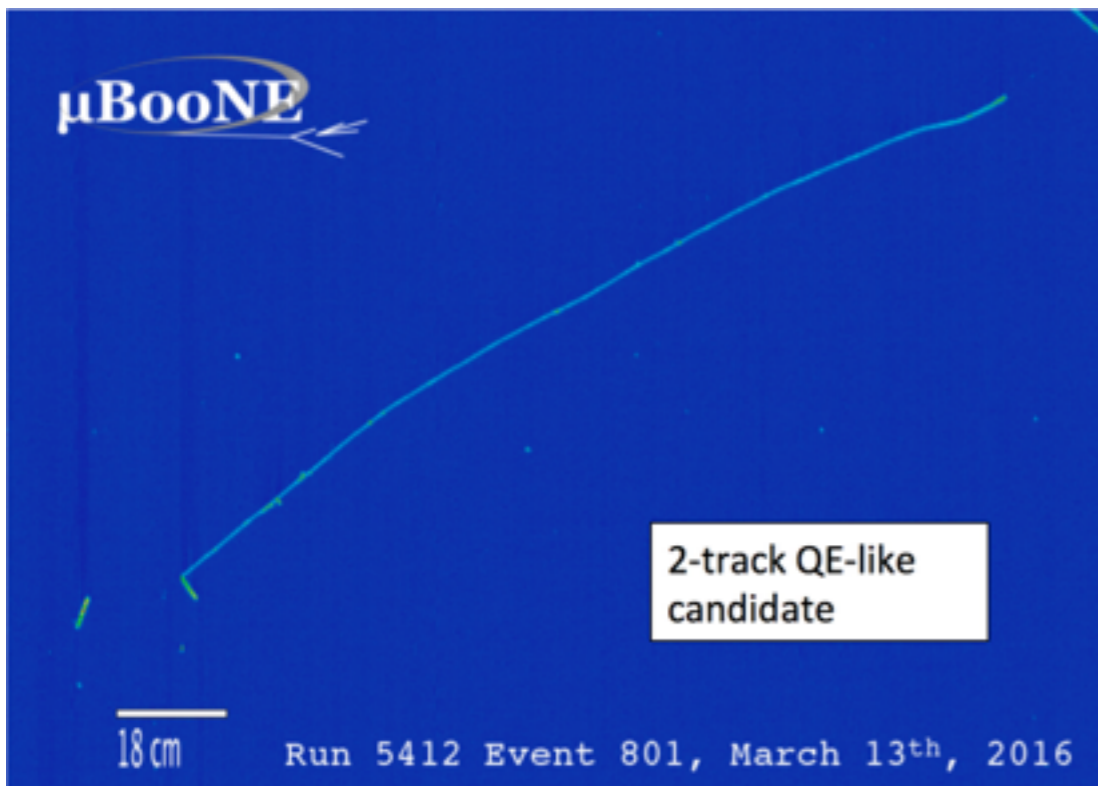


Intermediate Water-Cherenkov detector
Map detector response using multiple off-axis angles

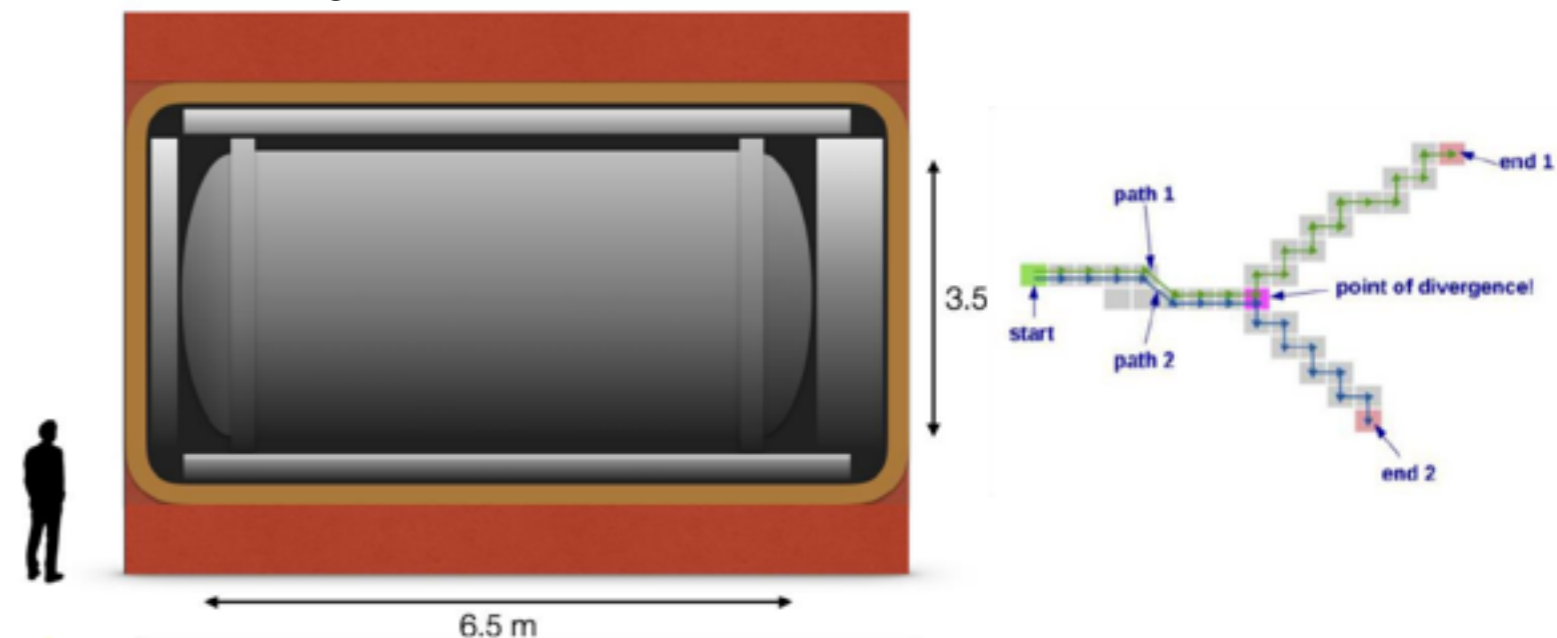
Near Detector Development

Several Argon TPC experiments
Natural ND candidates for DUNE

Precisely image the neutrino interaction vertex
(better constraints on neutrino-nucleus
interaction models → better energy
measurement)



DUNE High Pressure Gaseous TPC ND



Ultra-low thresholds with gaseous TPC

Summary

Statistical precision promised by future high beam power and high mass experiments place high demands on the systematic uncertainties that experiments must reach

T2K and NOvA have reported systematics uncertainties in the range $\sim 3 - 10\%$ level

Reductions are needed today to make best use of the increasing statistical precision in the T2K and NOvA disappearance measurements

Improved flux determination, ν -nucleus interaction modelling and understanding of detector response will all play a role



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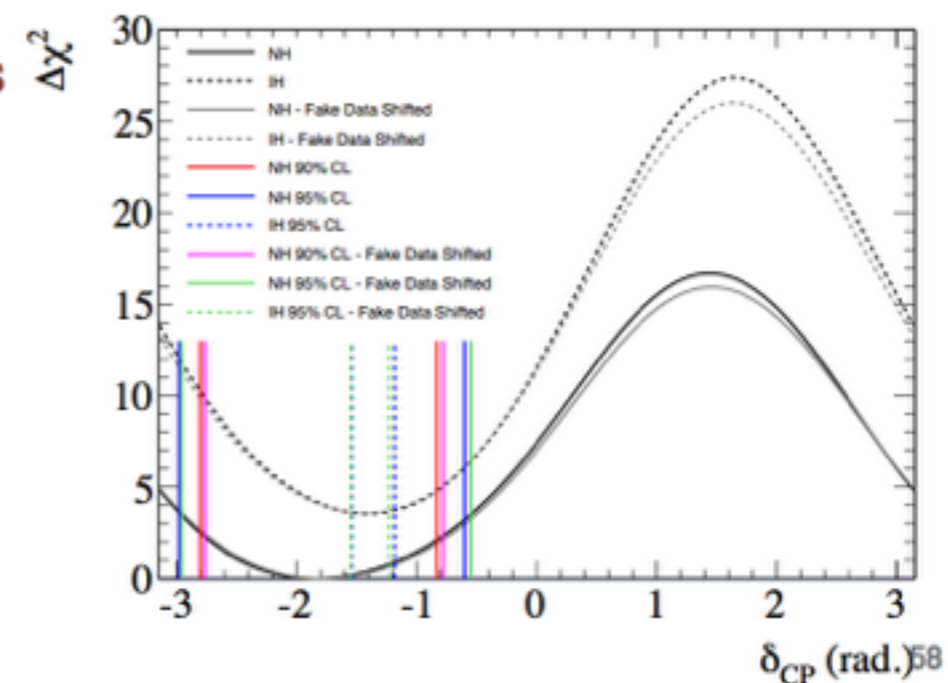
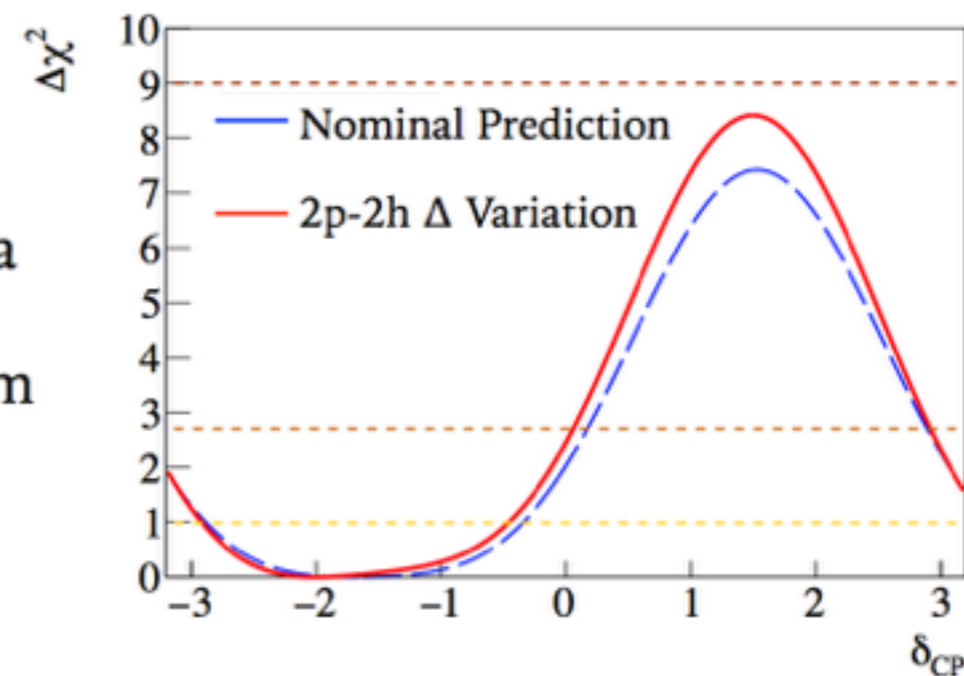
David Hadley
21st December 2017
Prospects in Neutrino Physics, NuPhys2017

Fake Data Studies

IMPACT ON CP PHASE



- Consider how changes to the $\Delta\chi^2$ impact intervals calculated from data
 - Shift $\Delta\chi^2$ observed in data (bottom plot) by difference observed in systematic study (top plot)
- **Maximum shift in the NH 2σ confidence interval mid-point was 1.7%**
- **Maximum change to the NH 2σ confidence interval was 2.3%**
- **Impact on δ_{CP} intervals is small**

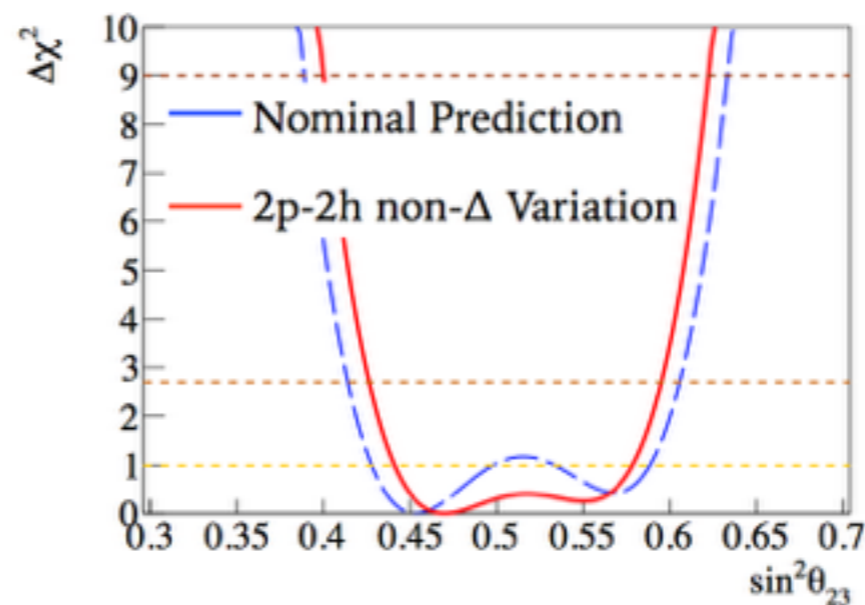
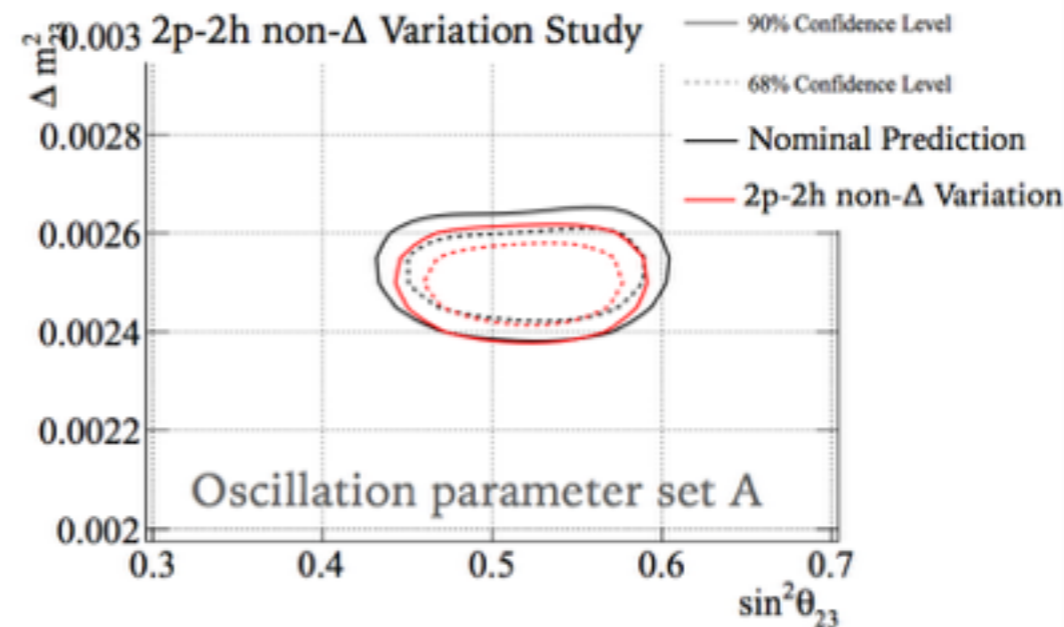


Fake Data Studies

IMPACT ON ATMOSPHERIC PARAMETERS



- ▶ In this study, Δm^2_{32} is biased to lower values
- ▶ $\sin^2\theta_{23}$ is biased towards maximal disappearance
 - ▶ Leads to narrower contour than fit to nominal prediction
- ▶ Shift towards maximal also seen in 1-D contour for oscillation parameter set B (bottom)



Fake Data Studies

ND280 DATA-DRIVEN VARIATION



➤ Take excess of data over prediction prior to ND280 fitting

➤ Assign excess to 1 of 3 types of interactions:

➤ CCQE

➤ 2p-2h Δ -enhanced

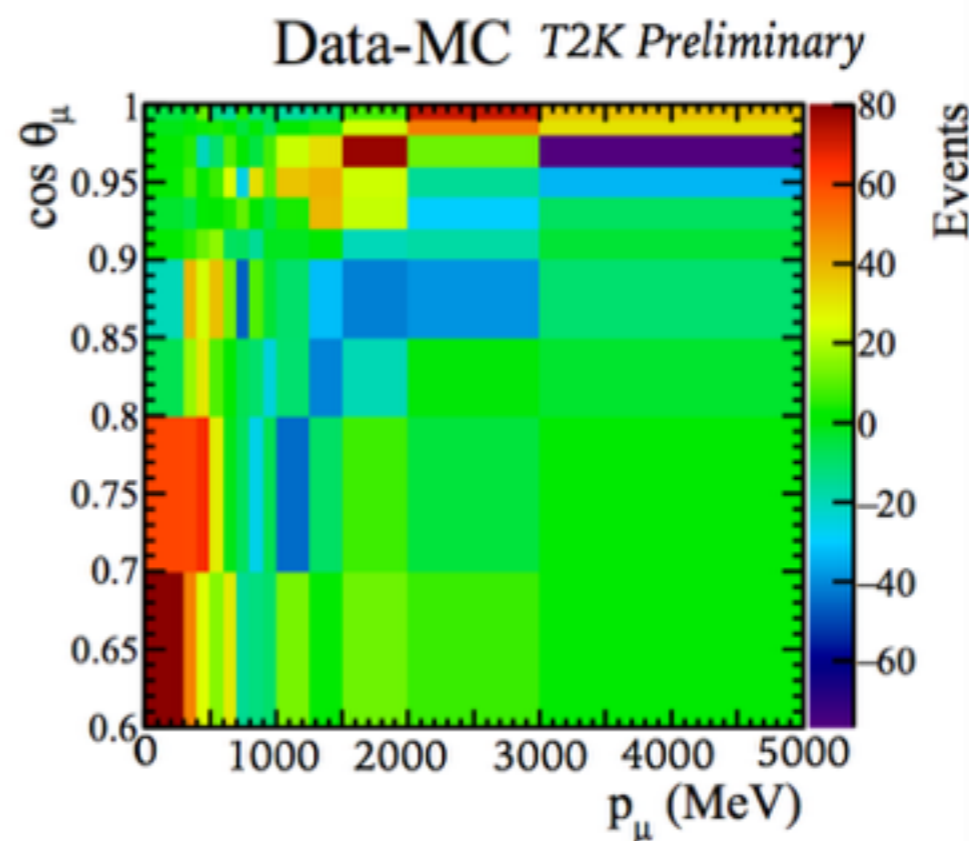
➤ 2p-2h non- Δ -enhanced

➤ Apply modeled excess to predict rates ND280 and SK

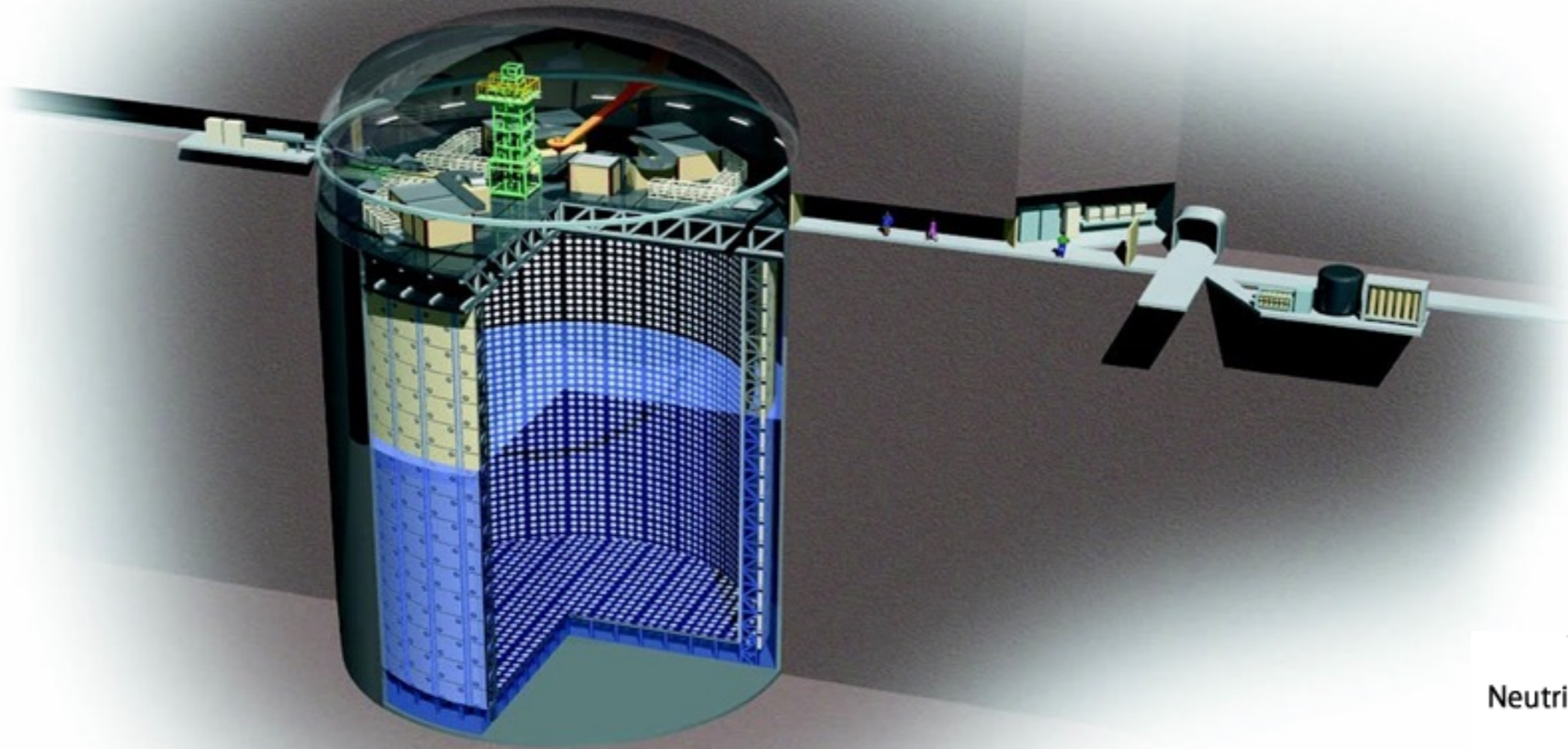
➤ Run fits

➤ **Effect seen on $\sin^2\theta_{23}$ and Δm^2_{32}**

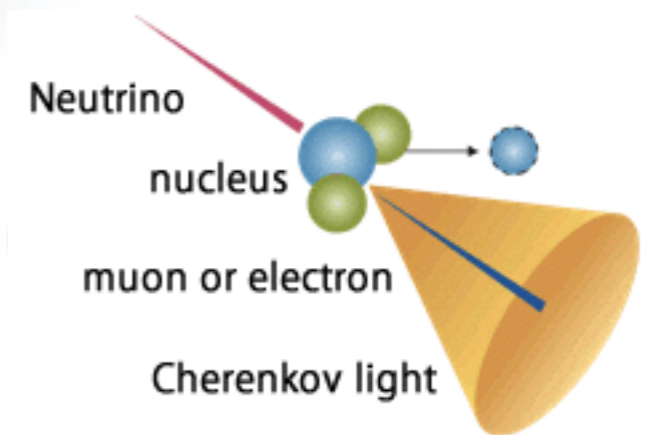
➤ **No significant impact on the measured intervals for δ_{cp}**



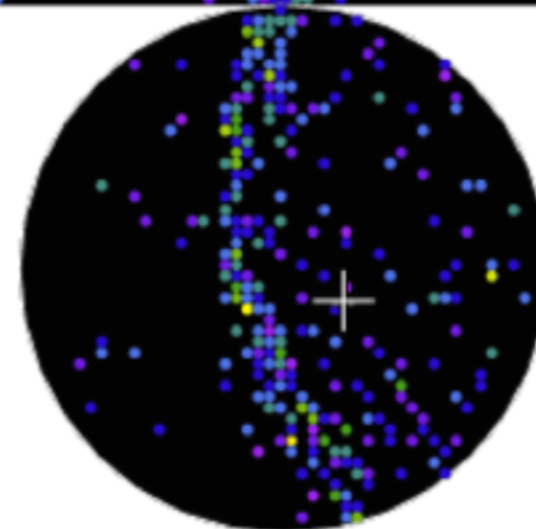
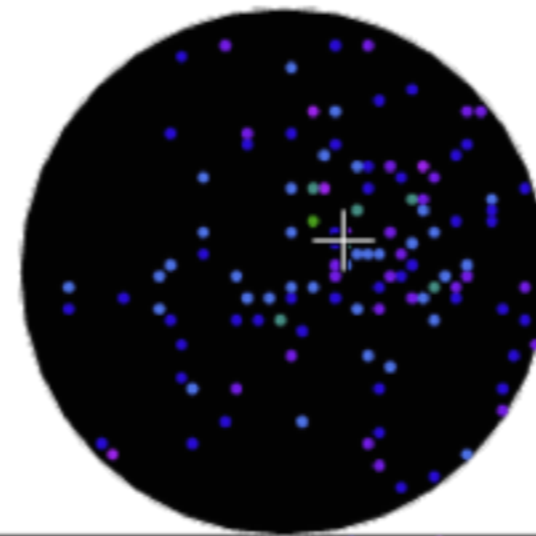
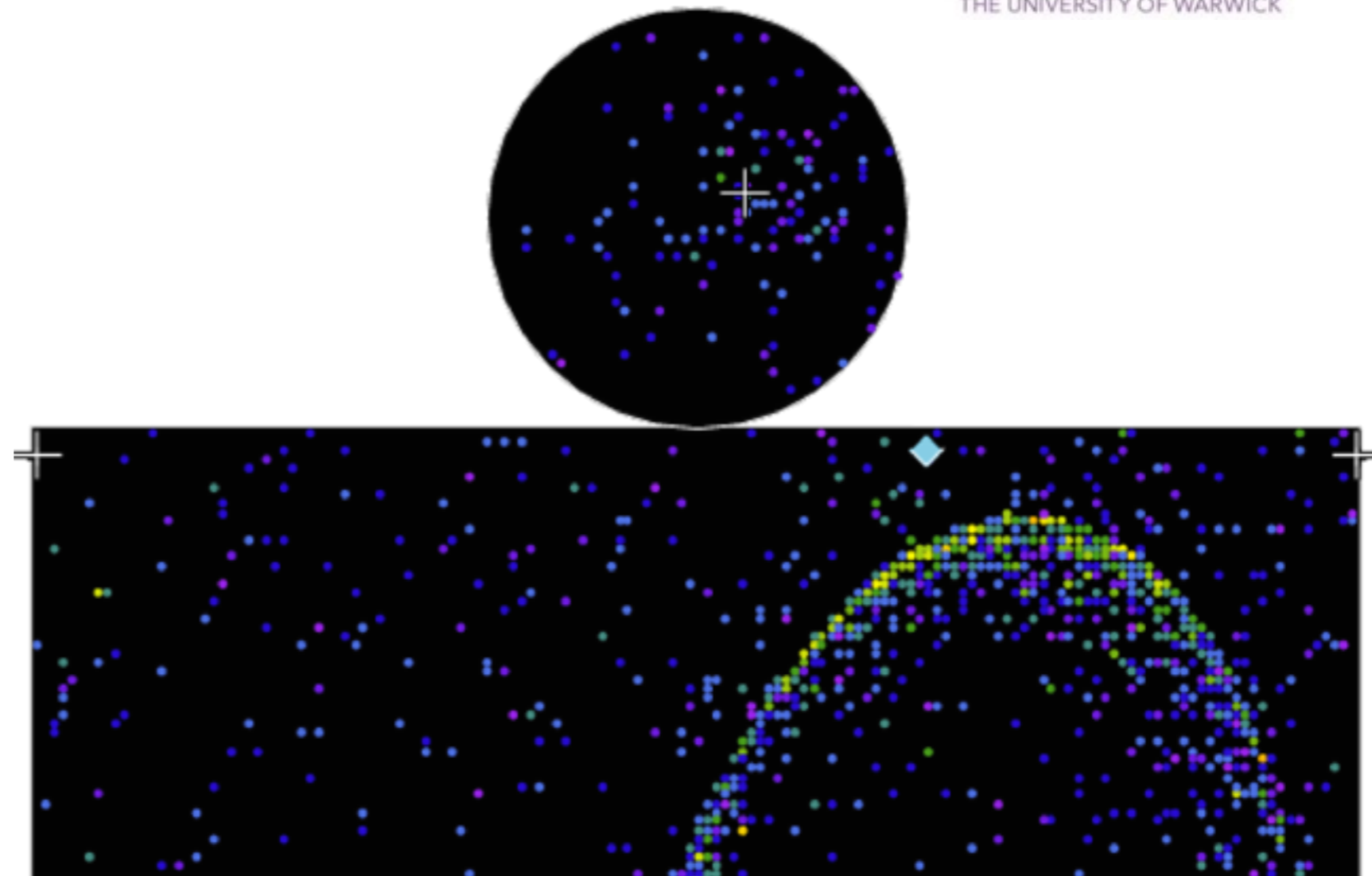
Super-Kamiokande



Water Cherenkov Detector
>22.5 kt fiducial mass



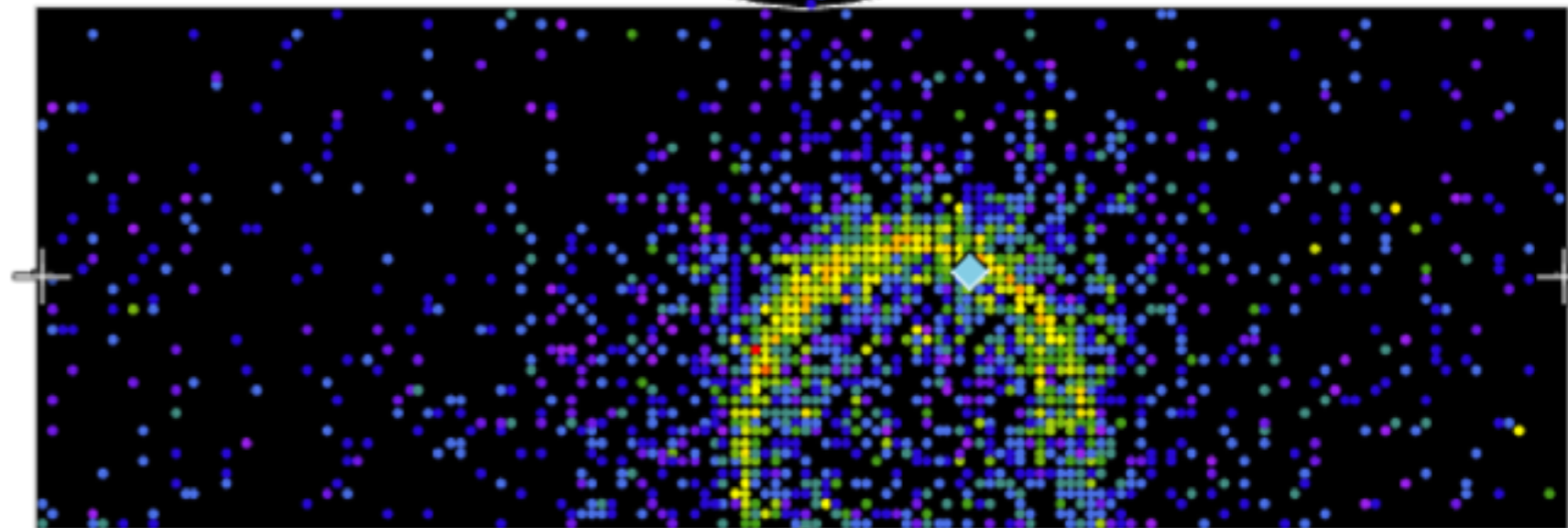
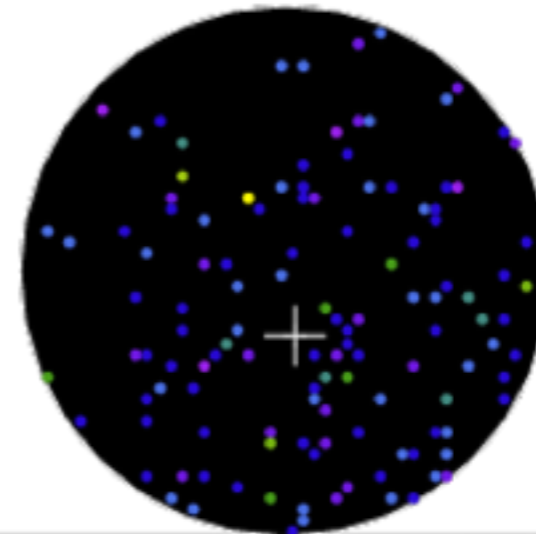
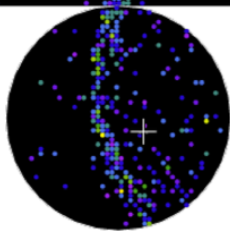
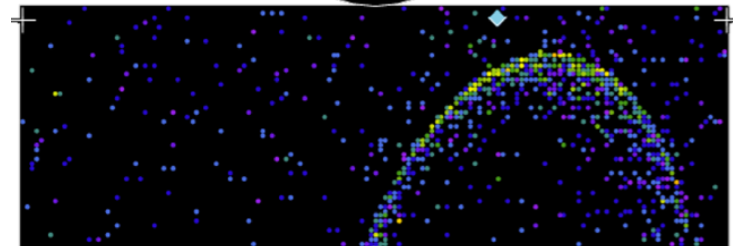
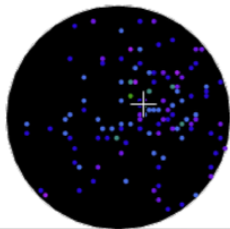
Water Cherenkov Technique



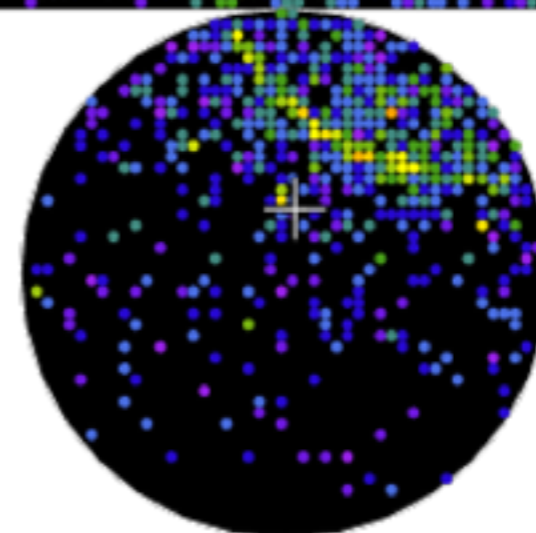
Muon

Water Cherenkov Technique

Muon

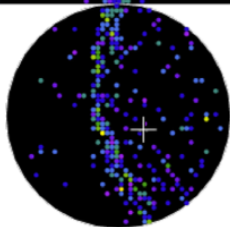
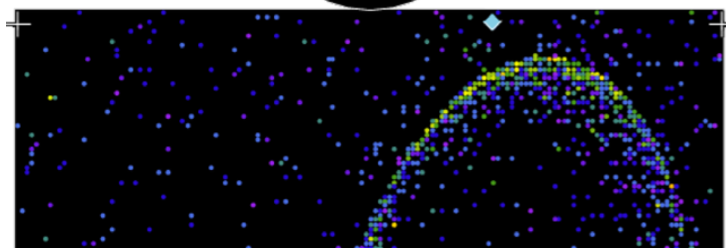
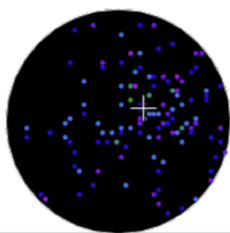


Electron

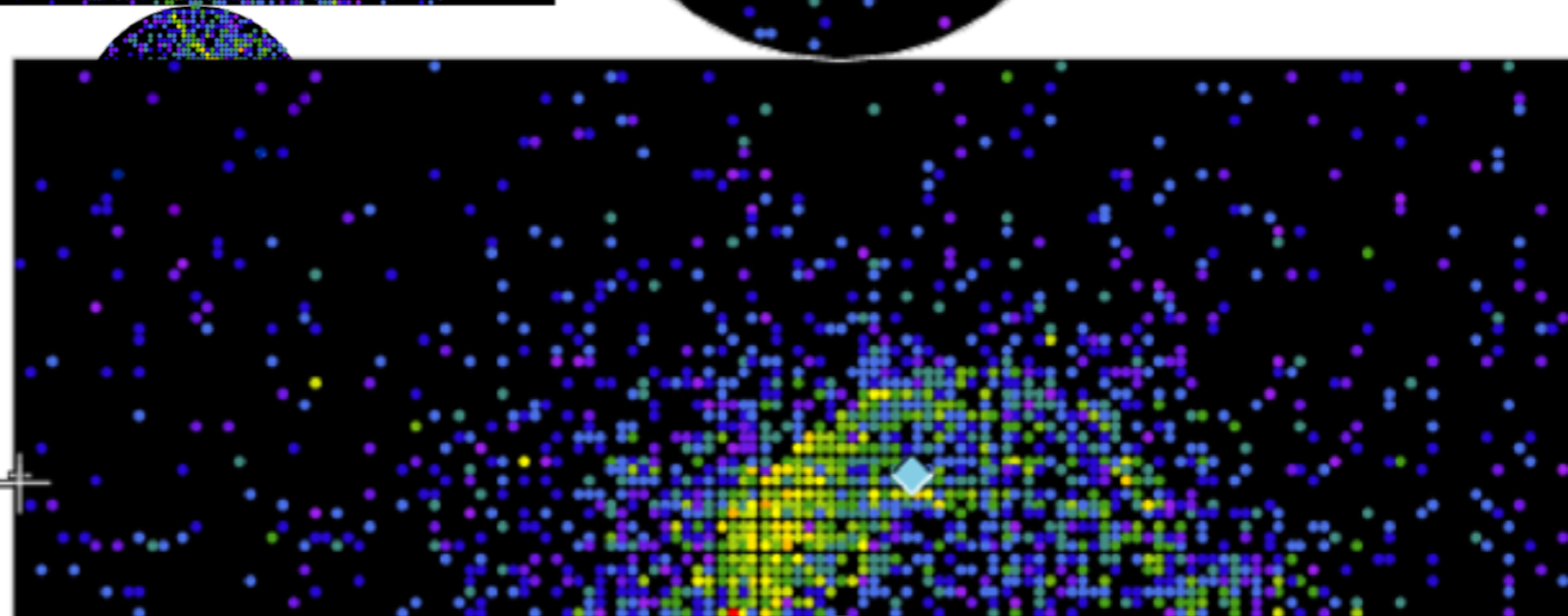
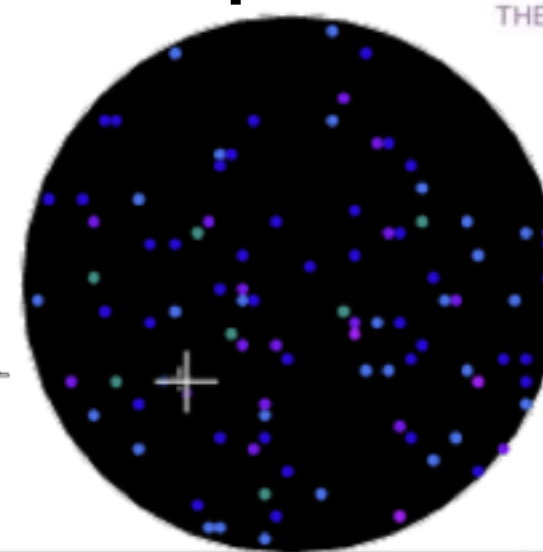
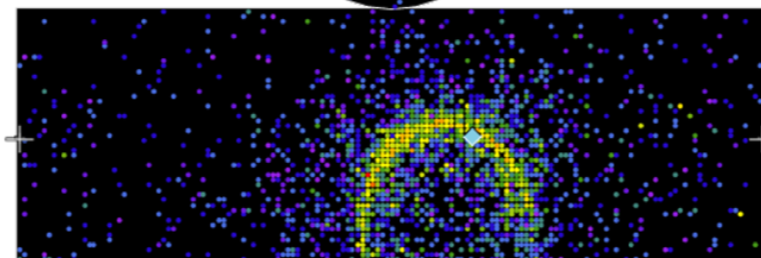
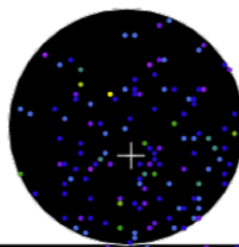


Water Cherenkov Technique

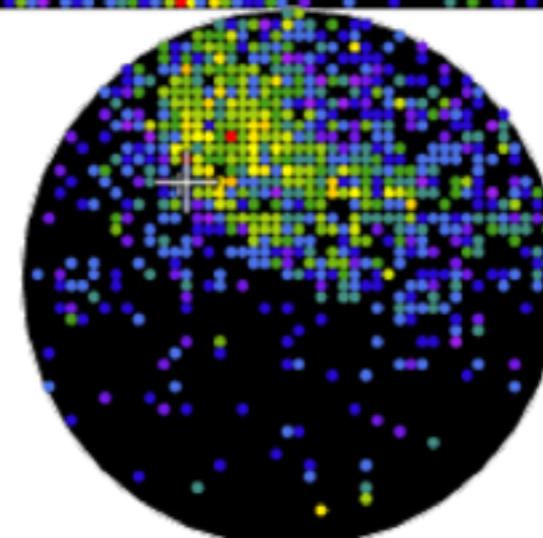
Muon



Electron

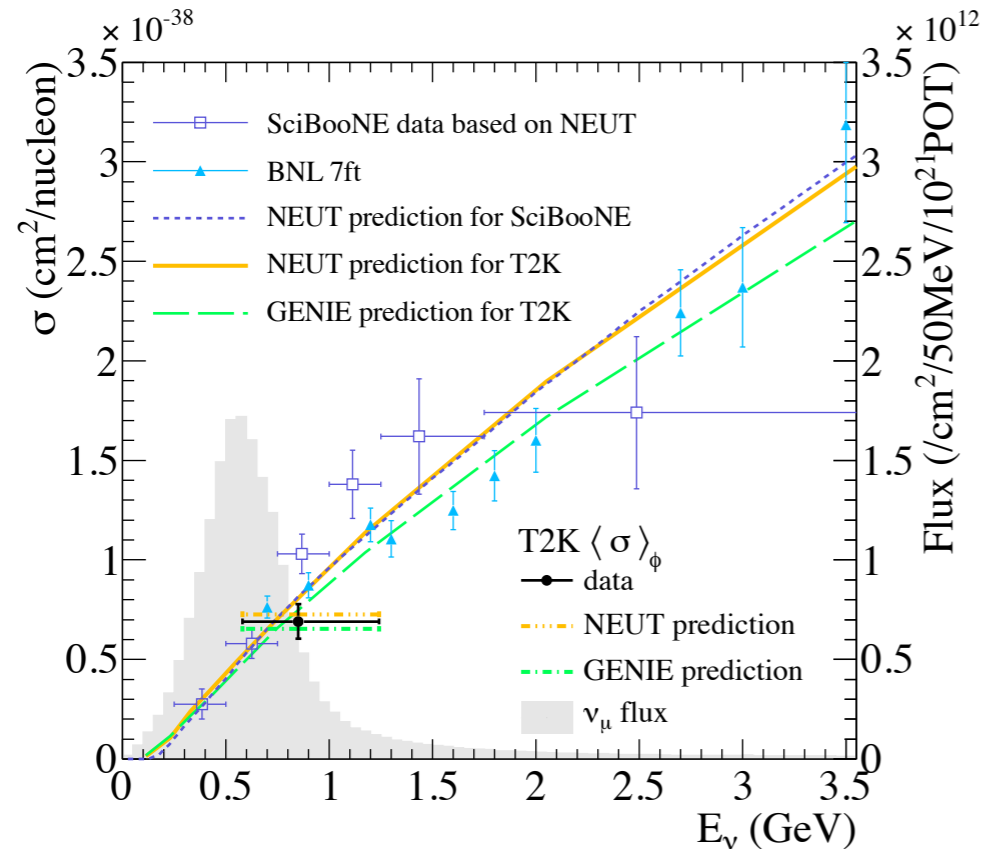


Neutral Pion

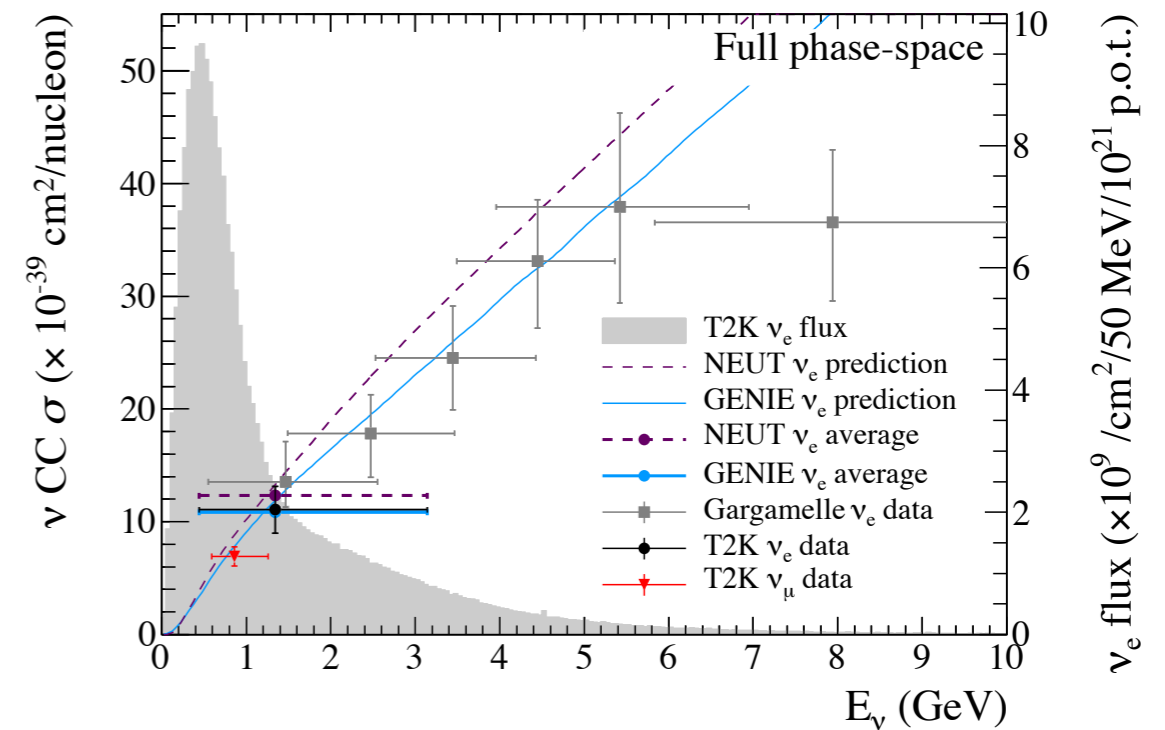


ND280 Flux

Phys. Rev. D 87, 092003 (2013)



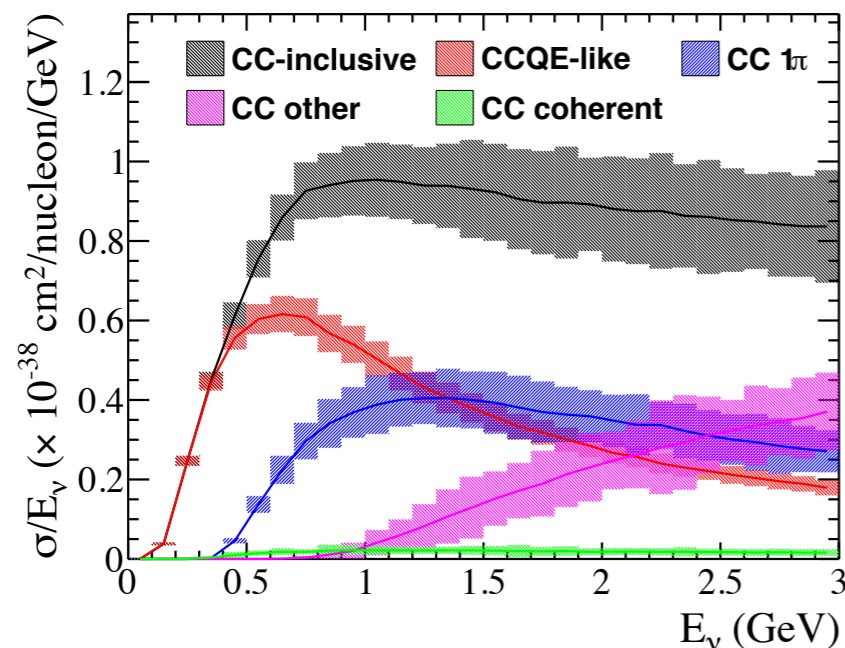
Phys. Rev. Lett. 113, 241803 (2014)



In neutrino-mode

ν_μ : $\langle E \rangle = 0.85$ GeV, ($\sim 90\%$)

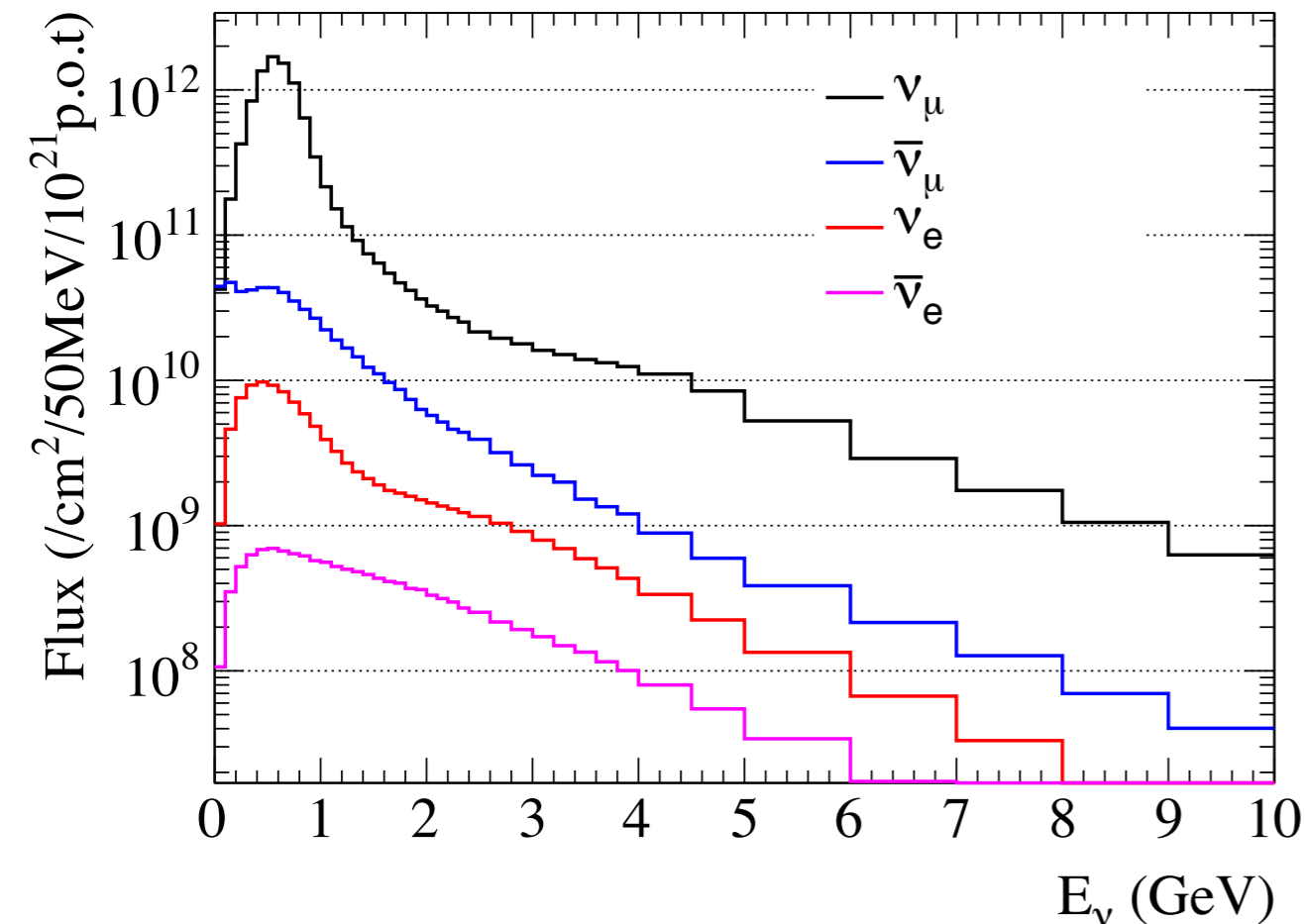
ν_e : $\langle E \rangle = 1.3$ GeV, ($\sim 1\%$)



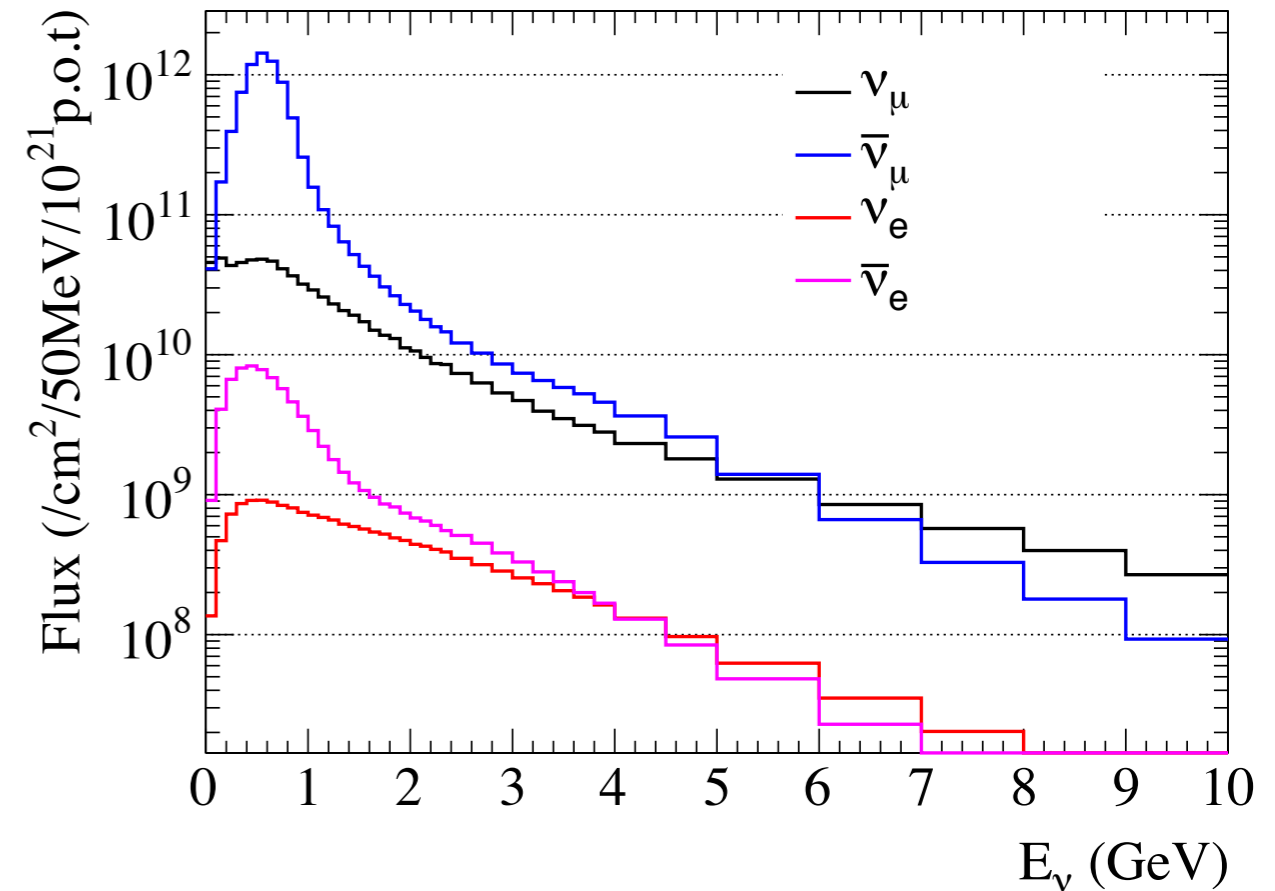
Dominant Reaction: CCQE
Single Pion Production

Flux at ND280

Neutrino Mode Flux at ND280



Antineutrino Mode Flux at ND280



In neutrino-mode

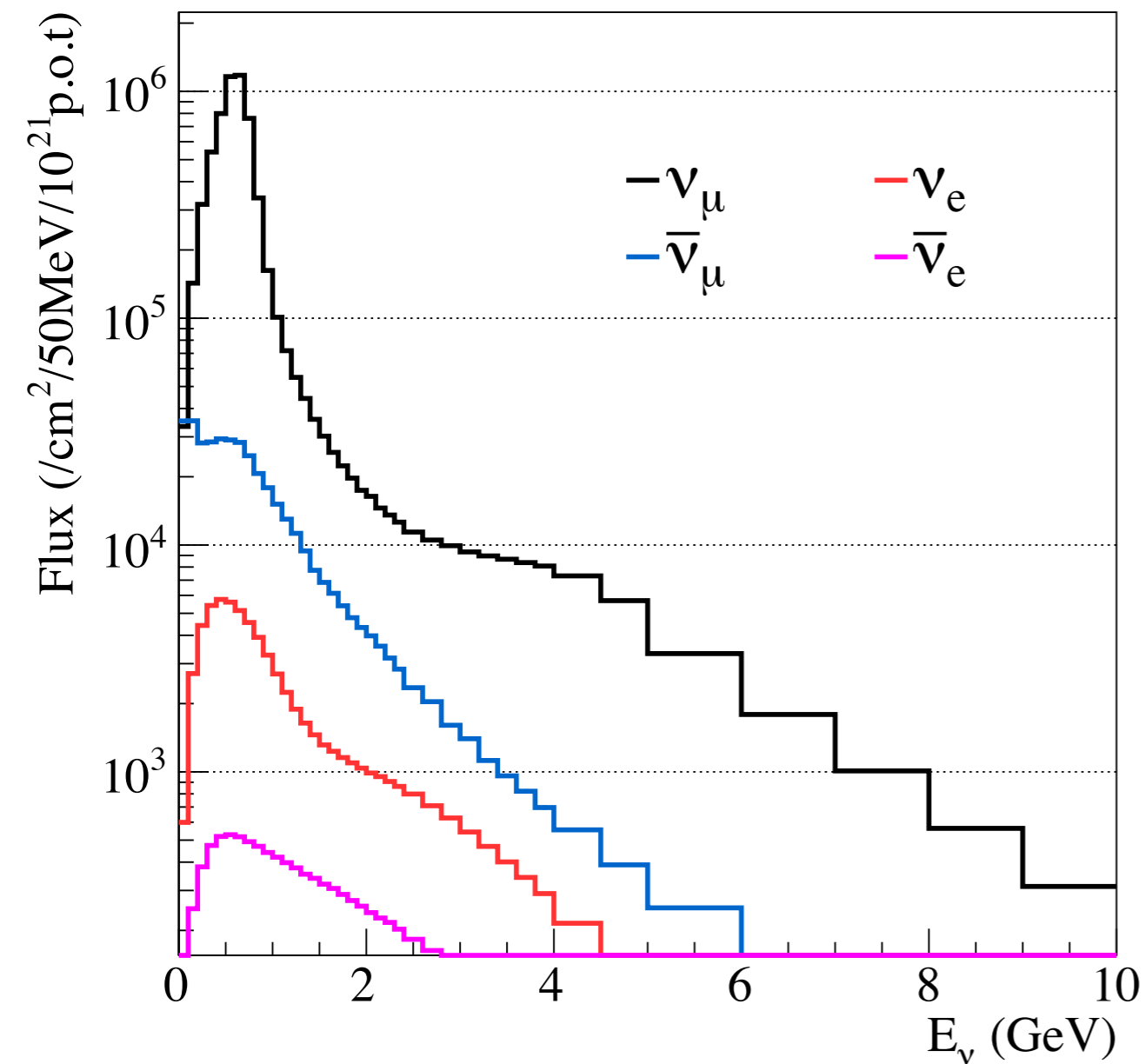
ν_μ : $\langle E \rangle = 0.85$ GeV, ($\sim 90\%$)

ν_e : $\langle E \rangle = 1.3$ GeV, ($\sim 1\%$)

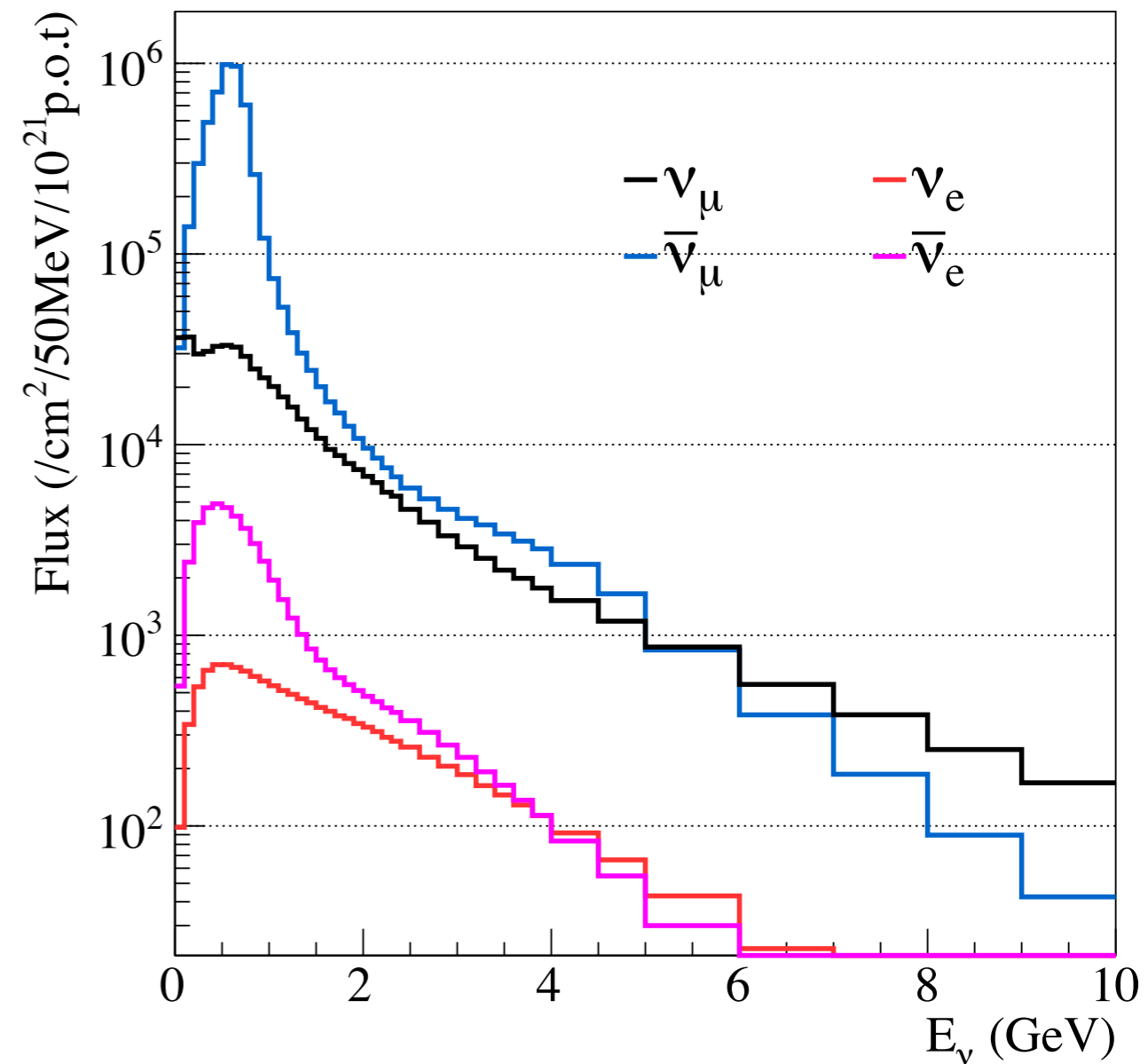
Dominant Reaction: CCQE
Single Pion Production

Flux at Super-K

Neutrino Mode Flux at SK

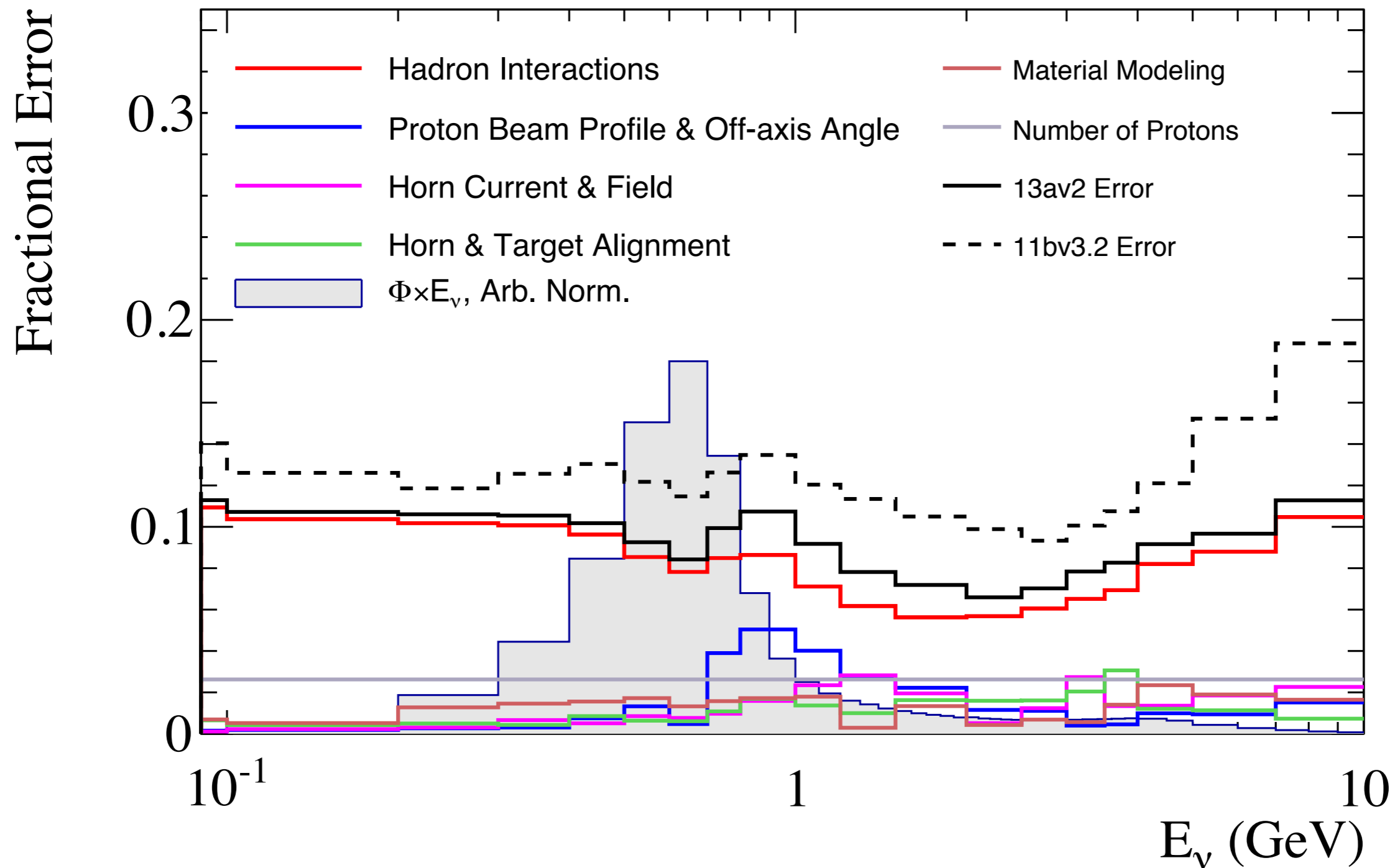


Antineutrino Mode Flux at SK



Flux Uncertainty

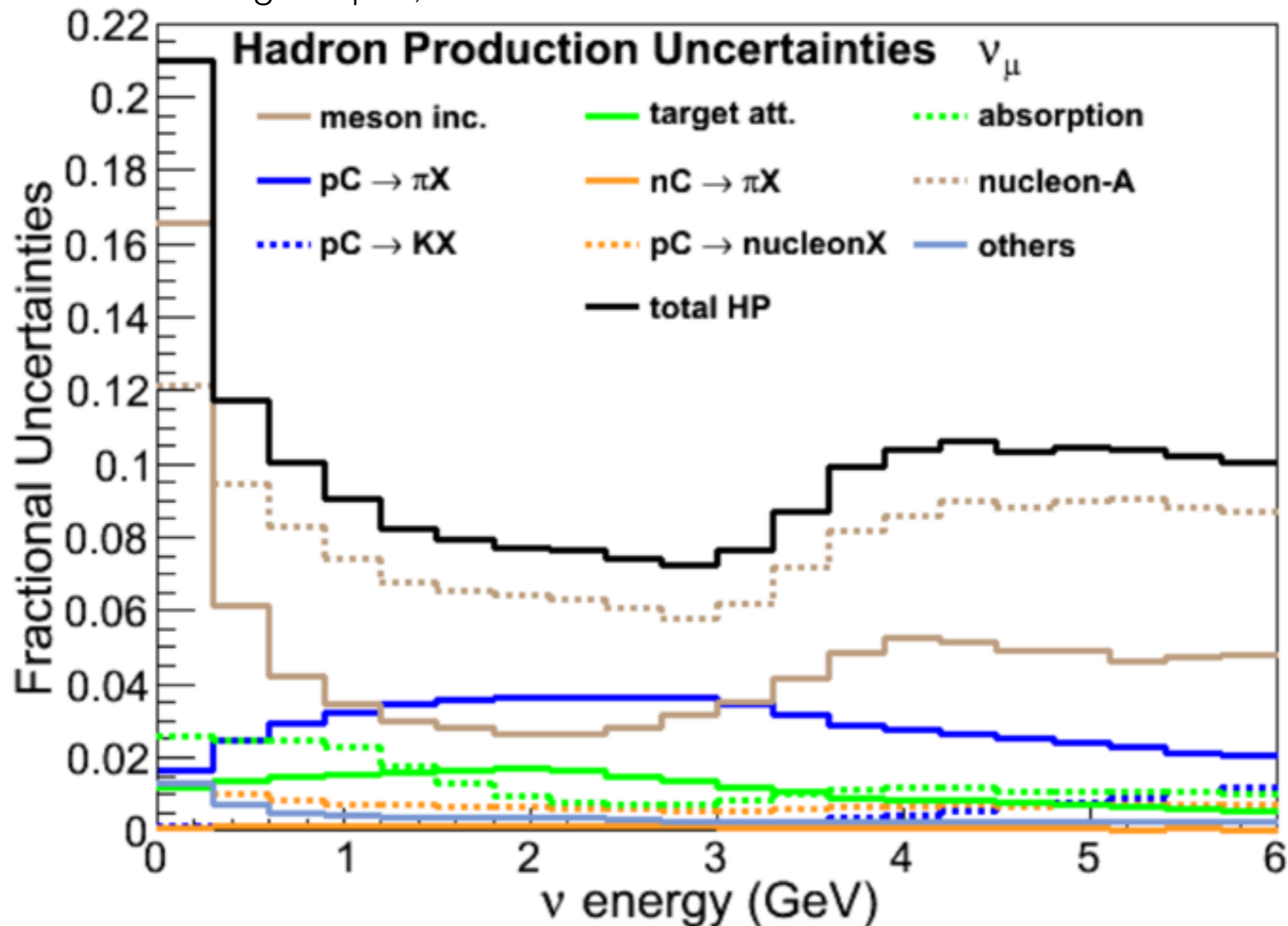
SK: Neutrino Mode, ν_μ



Flux Uncertainty

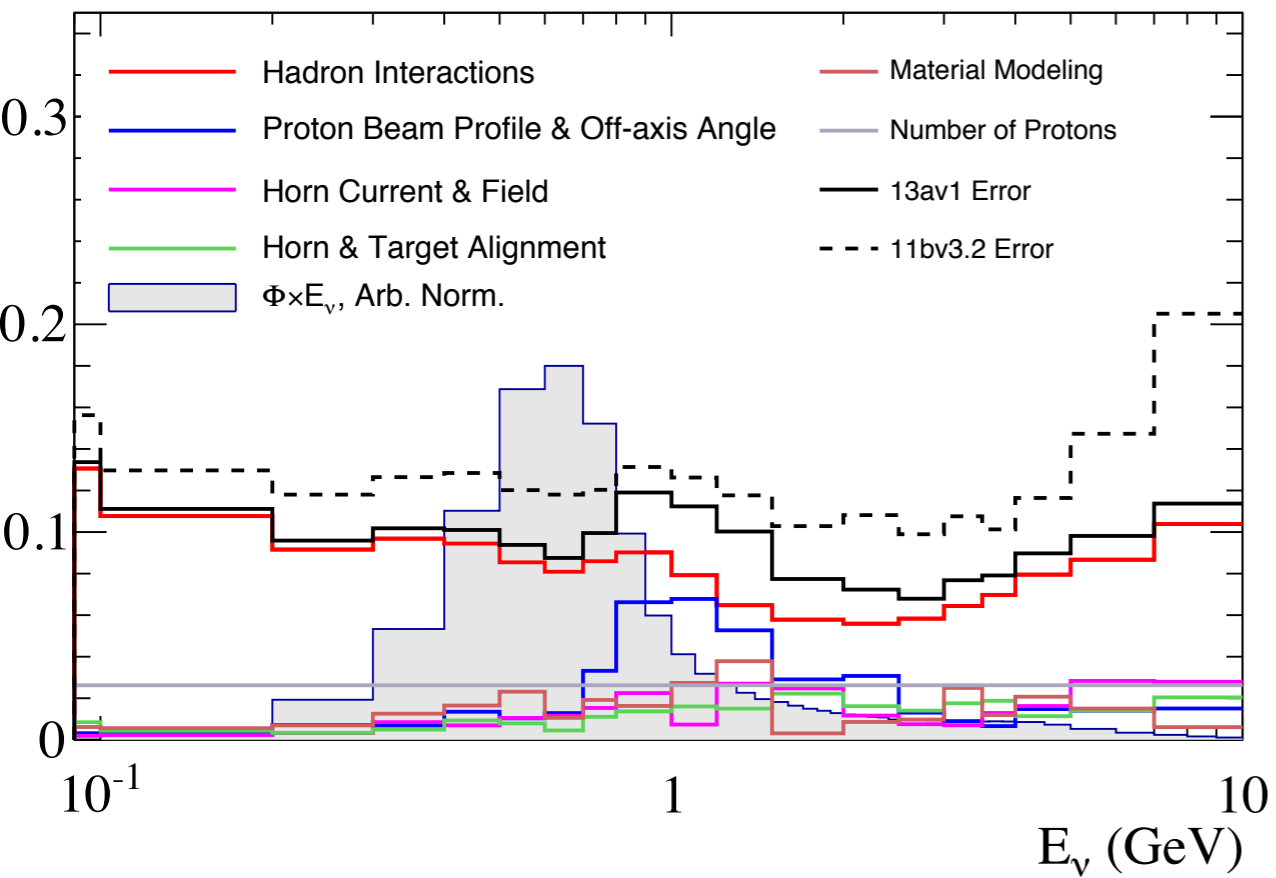
L Aliaga Soplin, NuInt 2017

NOvA Simulation

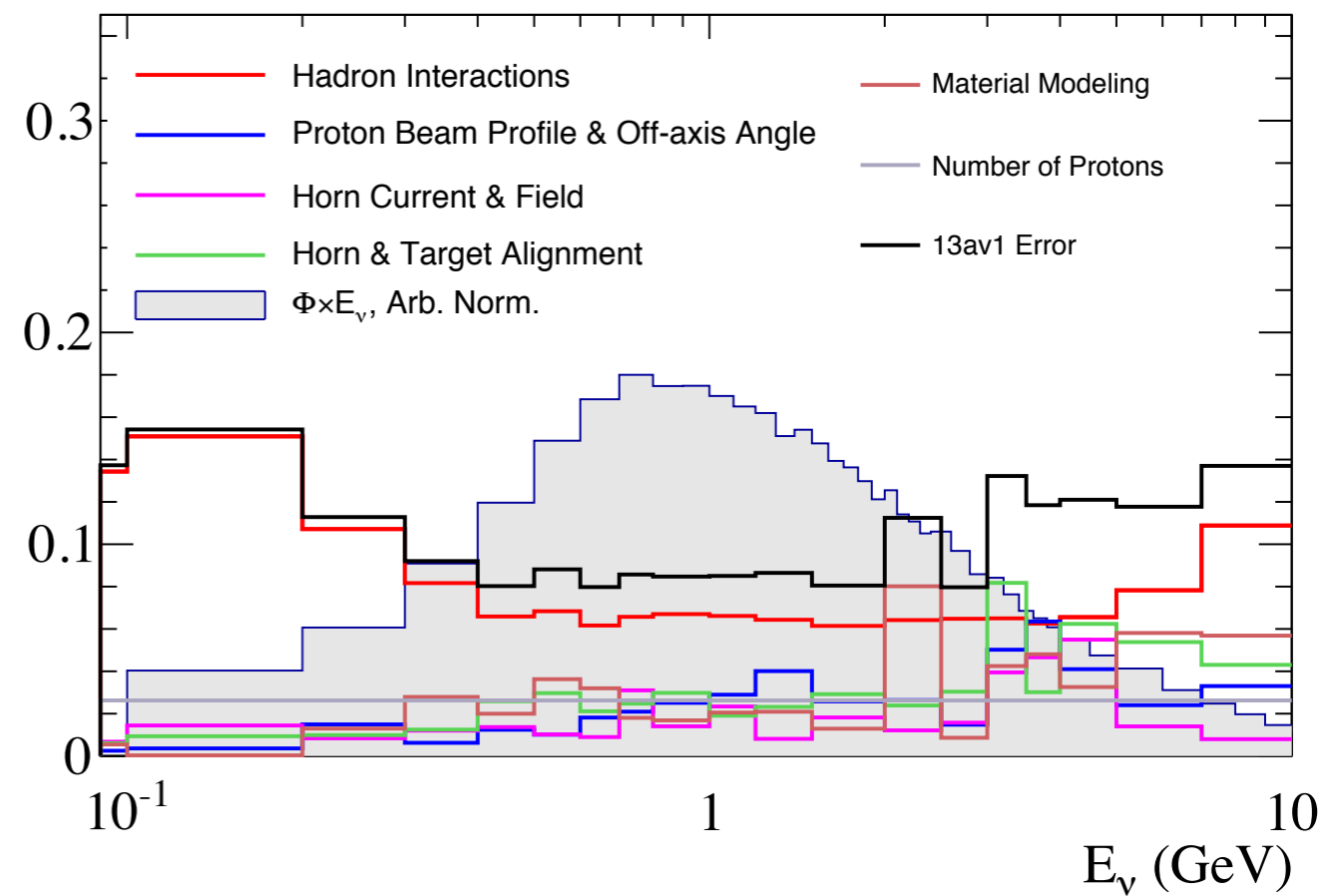


Flux at ND280

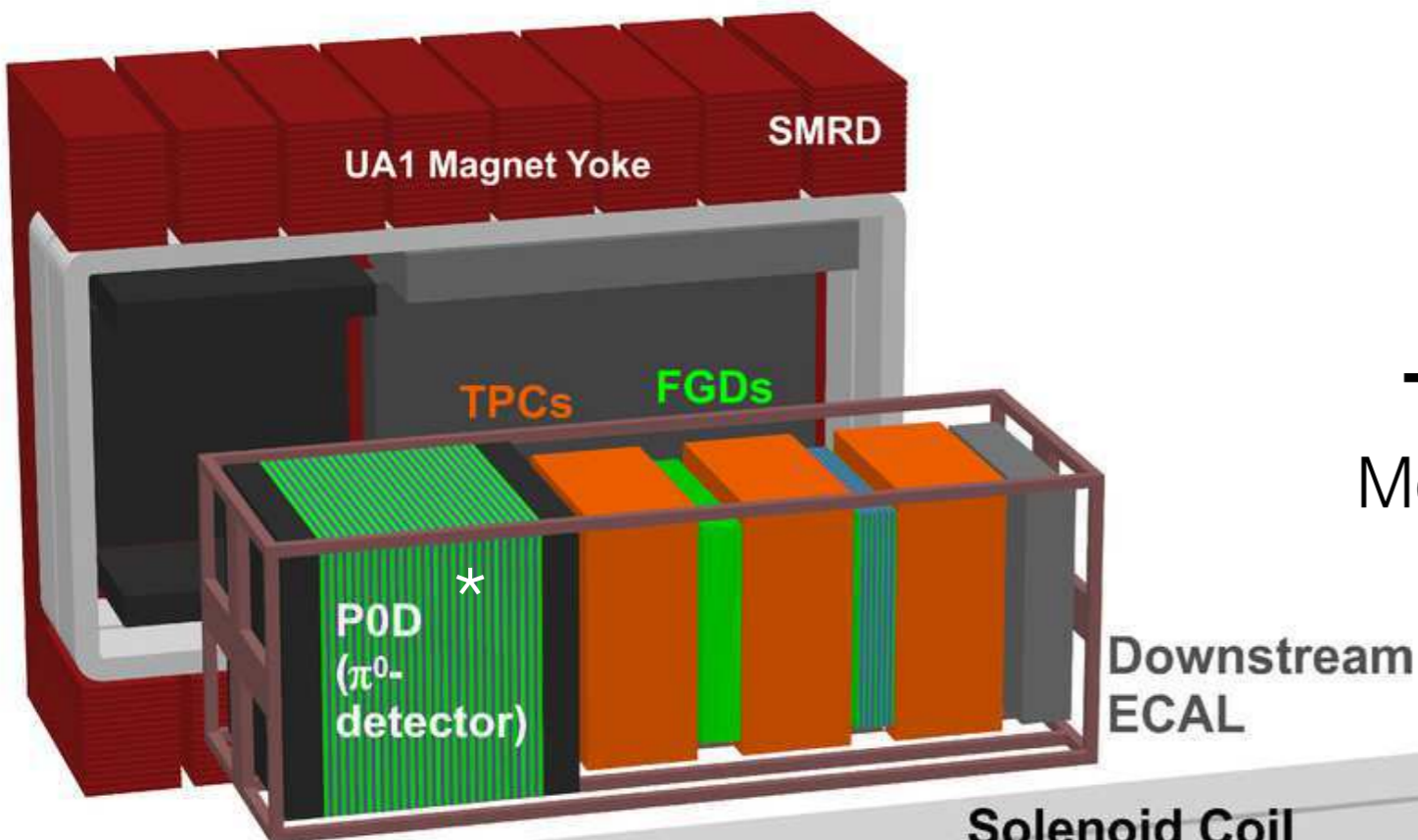
ND280: Neutrino Mode, ν_μ



ND280: Antineutrino Mode, $\bar{\nu}_\mu$

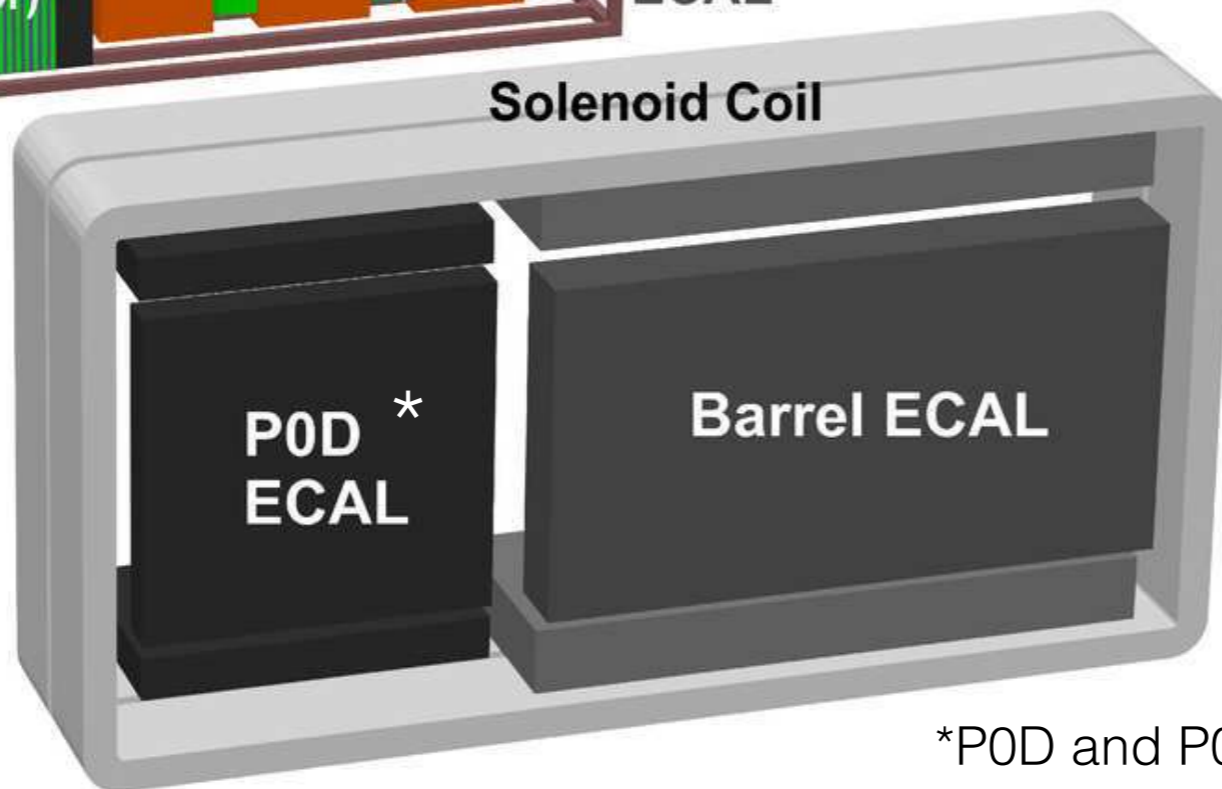
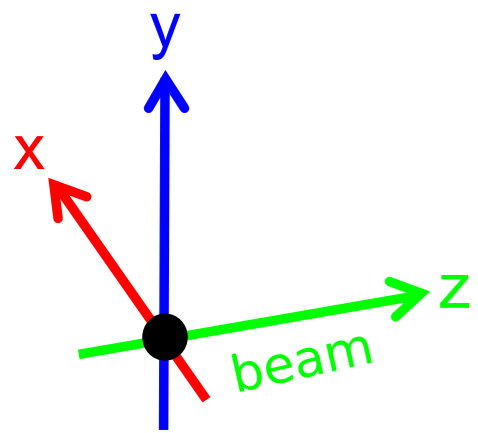


ND280 Detector



Fine Grained Detectors (FGD)
Carbon and Oxygen Target Mass,
Vertex reconstruction

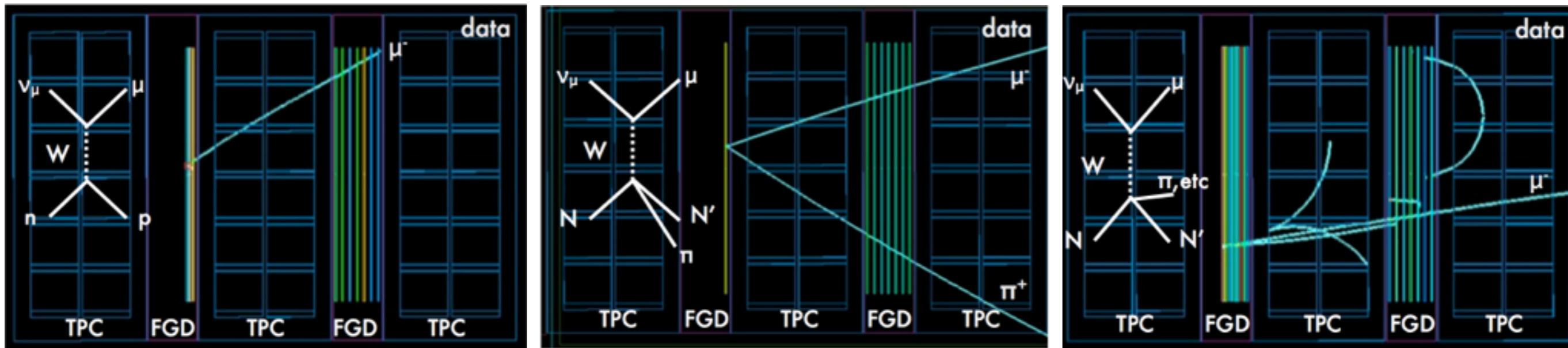
Time Projection Chambers (TPC)
Momentum and Charge Measurement
Particle ID



EM Calorimeters
Neutral Particle Reconstruction
Additional PID and
energy measurement
Tag entering backgrounds

*P0D and P0D ECAL detectors not be discussed here.
See arXiv:1111.5030 and arXiv:1308.3445 for information on these detectors.

ND280 Input to T2K Oscillation Analysis

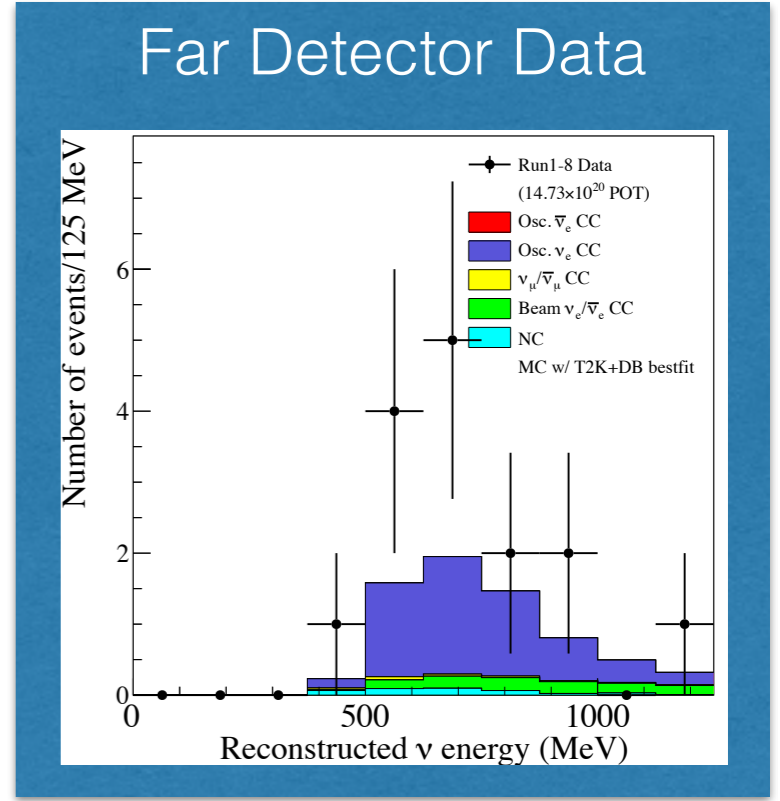
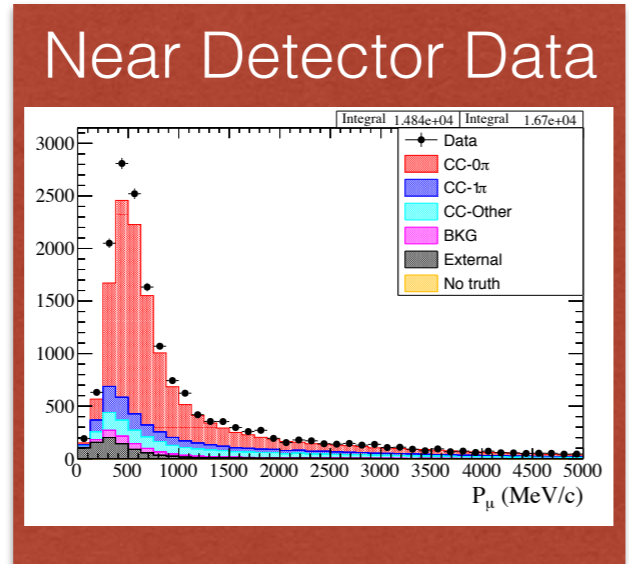


ND280 data split based on reconstructed topology enhanced in different interaction types

Fit flux + interaction model and propagate to far detector

As statistics increase and analysis becomes more sophisticated incorporate more channels

T2K Analysis Strategy



Theoretical Predictions

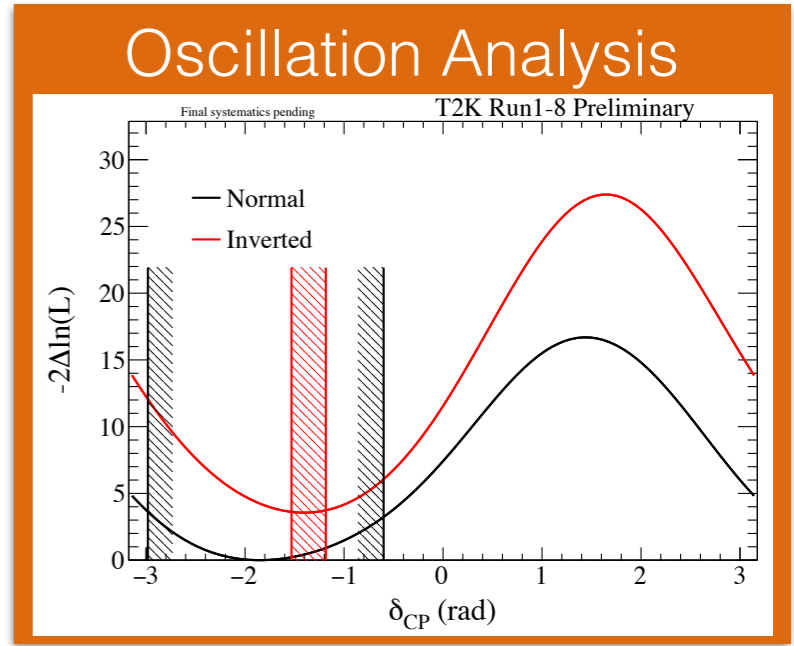
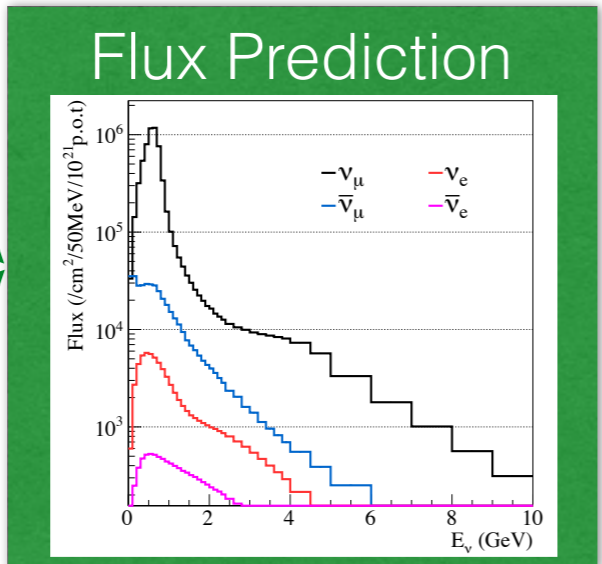
External Cross Section Measurements

ν -Nucleus Interaction Model

Updated Flux + ν -Nucleus Interaction Model

NA61/SHINE

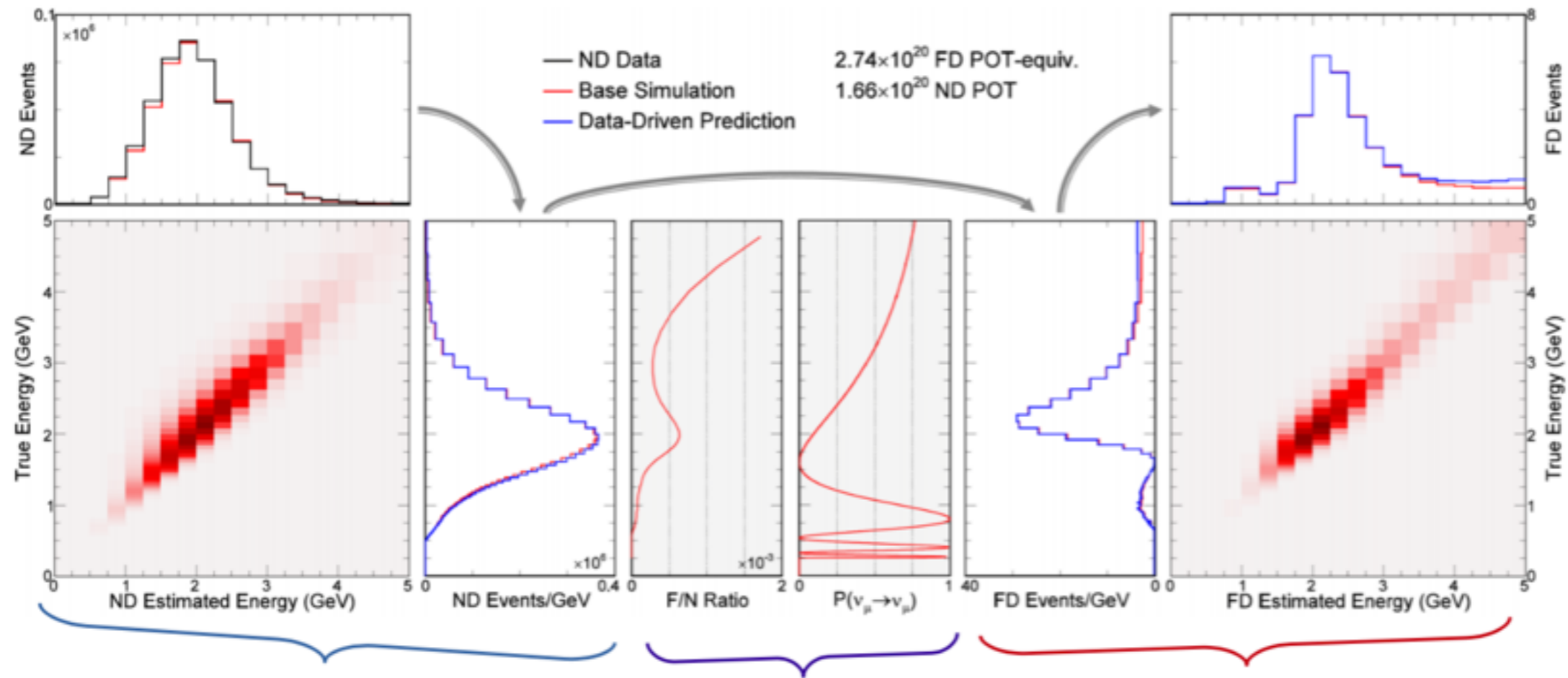
INGRID and Beam Monitors



NOvA Analysis Strategy

To produce a data-driven prediction at FD, based on ND:

J. Wolcot, NuInt 2017



True energy distribution is corrected so that reconstructed data & MC agree at the ND...

...modified true energy distribution is propagated through predicted geometric beam dispersion & acceptance ratio, oscillations...

... and "extrapolated" reconstructed energy distribution computed to compare to data